



SKA Computing Challenges

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Computing for SKA Colloquium
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The Square Kilometre Array

- Will be World's largest and most powerful radio telescope
- Actually four types of telescope arrays:
 - Low - sparse aperture array with 1024 stations, each with 256 dipole antennas, later also dense aperture array
 - Mid – 254 single pixel dishes
 - Survey – 96 phased-array feed receivers each with 36 beams
- Cost €650M for SKA1
- Early estimates €2- €6billion for SKA2
- Preconstruction design for SKA1 commenced Nov 2013
- Two precursor arrays already being built:
 - MeerKAT in South Africa (Mid)
 - ASKAP in Australia (Survey)
- Actually an enormous engineering design and construction project



SKA₁ Key Computing-Data Generation



Mid:
254 dishes
x 2 pol

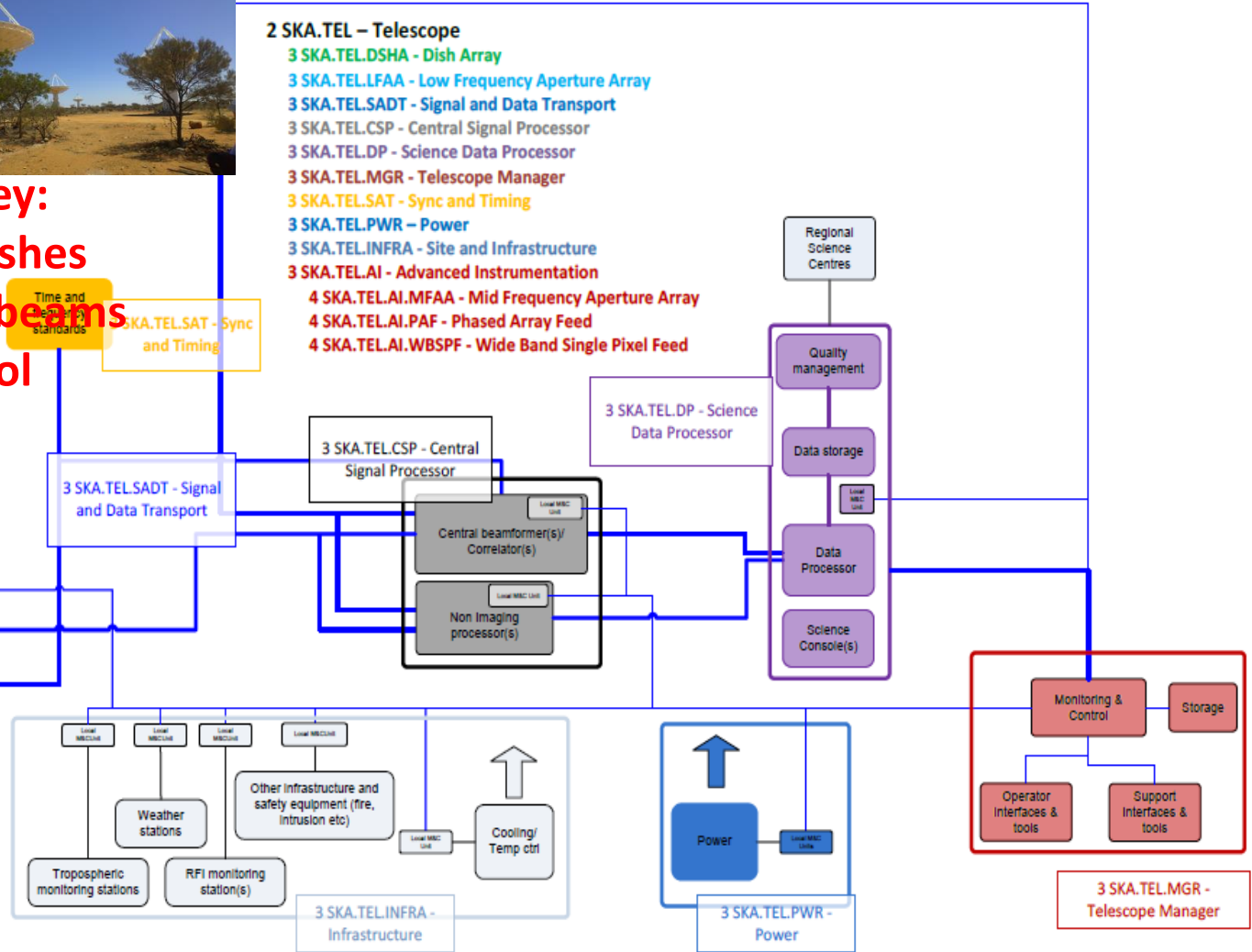
Survey:
96 dishes
x 36 beams
x 2 pol

Drawing:
Stage 1 WBS Prelim Block Diagram
Updated: 2011-12-09
Revision: B



Low:
1024 stations
x 256 antennas
x 2 pol

- 2 SKA.TEL – Telescope**
- 3 SKA.TEL.DSHA - Dish Array
 - 3 SKA.TEL.LFAA - Low Frequency Aperture Array
 - 3 SKA.TEL.SADT - Signal and Data Transport
 - 3 SKA.TEL.CSP - Central Signal Processor
 - 3 SKA.TEL.DP - Science Data Processor
 - 3 SKA.TEL.MGR - Telescope Manager
 - 3 SKA.TEL.SAT - Sync and Timing
 - 3 SKA.TEL.PWR – Power
 - 3 SKA.TEL.INFRA - Site and Infrastructure
 - 3 SKA.TEL.AI - Advanced Instrumentation
 - 4 SKA.TEL.AI.MFAA - Mid Frequency Aperture Array
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 - 4 SKA.TEL.AI.WBSPF - Wide Band Single Pixel Feed



3 SKA.TEL.MGR - Telescope Manager

SKA “Big Data”



The animation summarises the amount of data that would be generated by the current baseline design for SKA1.....assuming the science use cases that would maximise data rates for this design. Expect that, by construction time, the figures are likely to be reduced considerably.....although they will still be immense.

SKA₁ Key Computing-Beamforming



Survey Beamformer

3 SKA.TEL.DSHA - Dish Array

Survey Beamformer:
Each dish combines PAF signals into 36 beams

Time and frequency standards and Timing

Drawing: Stage1WBSPrelimBlockDiagram
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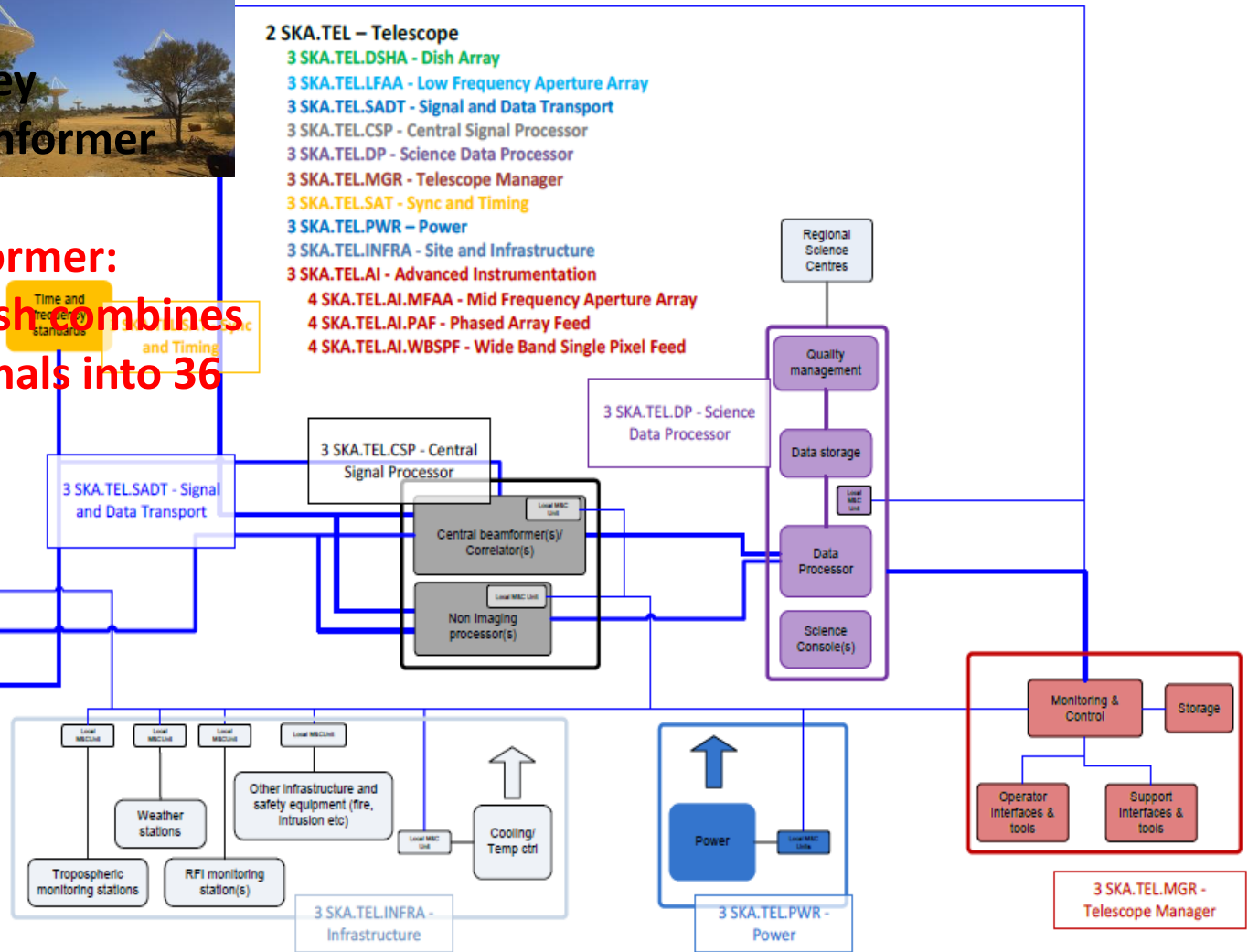


3 SKA.TEL.LFAA - Low Frequency Aperture Array

Low Beamformer:
Each station combines 256 antenna signals

2 SKA.TEL – Telescope

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SKA₁ Key Computing-Correlators



3 SKA.TEL.DSHA - Dish Array

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Stage1WBSPrelimBlockDiagram
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Time and frequency standards

3 SKA.TEL.SAT - Sync and Timing

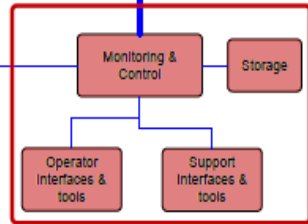
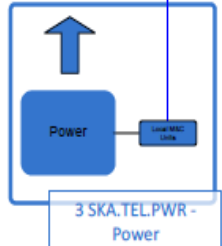
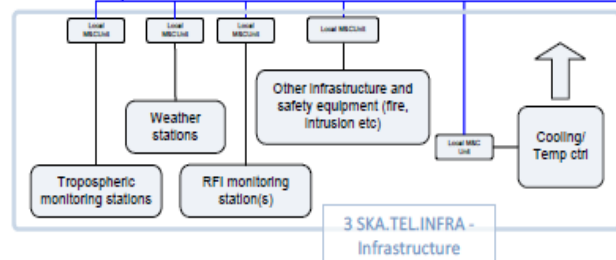
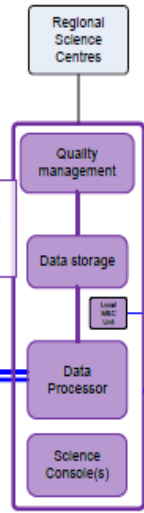
3 SKA.TEL.SADT - Signal and Data Transport

3 SKA.TEL.CSP - Central Signal processor

Low Correlator
Mid Correlator
Survey Correlator

Non Imaging processor(s)

3 SKA.TEL.DP - Science Data Processor



3 SKA.TEL.MGR - Telescope Manager

Correlator Overview

- Correlator combines signals together from each receiver so array acts like a single telescope
- FX correlator:
 - Each signal Fourier transformed
 - Each pair of transformed signals cross-correlated
 - Correlated products accumulated over brief time
 - Outputs are called visibilities
- Pairing requires interconnect between channelizer and correlator
 - To avoid enormous data duplication requires “centralised” correlator

Telescope	Data Input	Channelizer Compute	Correlator Compute	Max Data Output
Low	2338 GB/s	46.7 TMAC/s	2090 TMAC/s	7320 GB/s
Mid	580 GB/s	11.6 TMAC/s	2003 TMAC/s	3250 GB/s
Survey	7900 GB/s	158.0 TMAC/s	1313 TMAC/s	4670 GB/s

Correlator Challenges

- Exascale Big Data project:
 - Large Hadron Collider expected to produce about 15 PetaByte/year
 - Correlators will produce about 40 PetaByte/hour
- Unprecedented I/O challenge:
 - Each telescope array requires up to 8 TeraByte/s channelization and correlation and outputs up to 4 TeraByte/s (Cray Titan I/O about 1.4 TeraByte/s total)
 - Data throughput a major concern
- Green computing requirements:
 - Two telescope arrays are deep in Australian outback, no power grid so energy consumption a major concern

SKA₁ Key Computing-Pulsar Search



3 SKA.TEL.DSHA - Dish Array

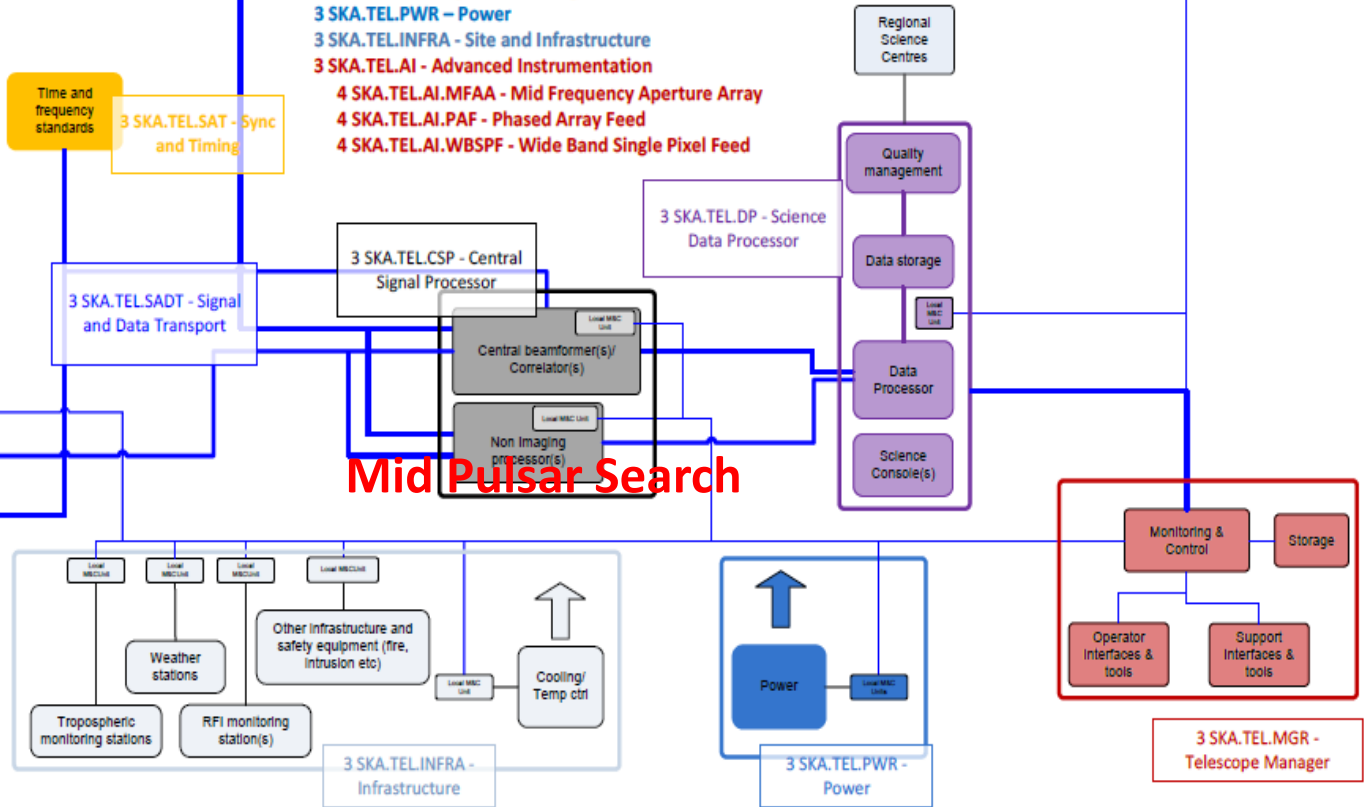
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3 SKA.TEL.LFAA - Low Frequency Aperture Array

2 SKA.TEL – Telescope

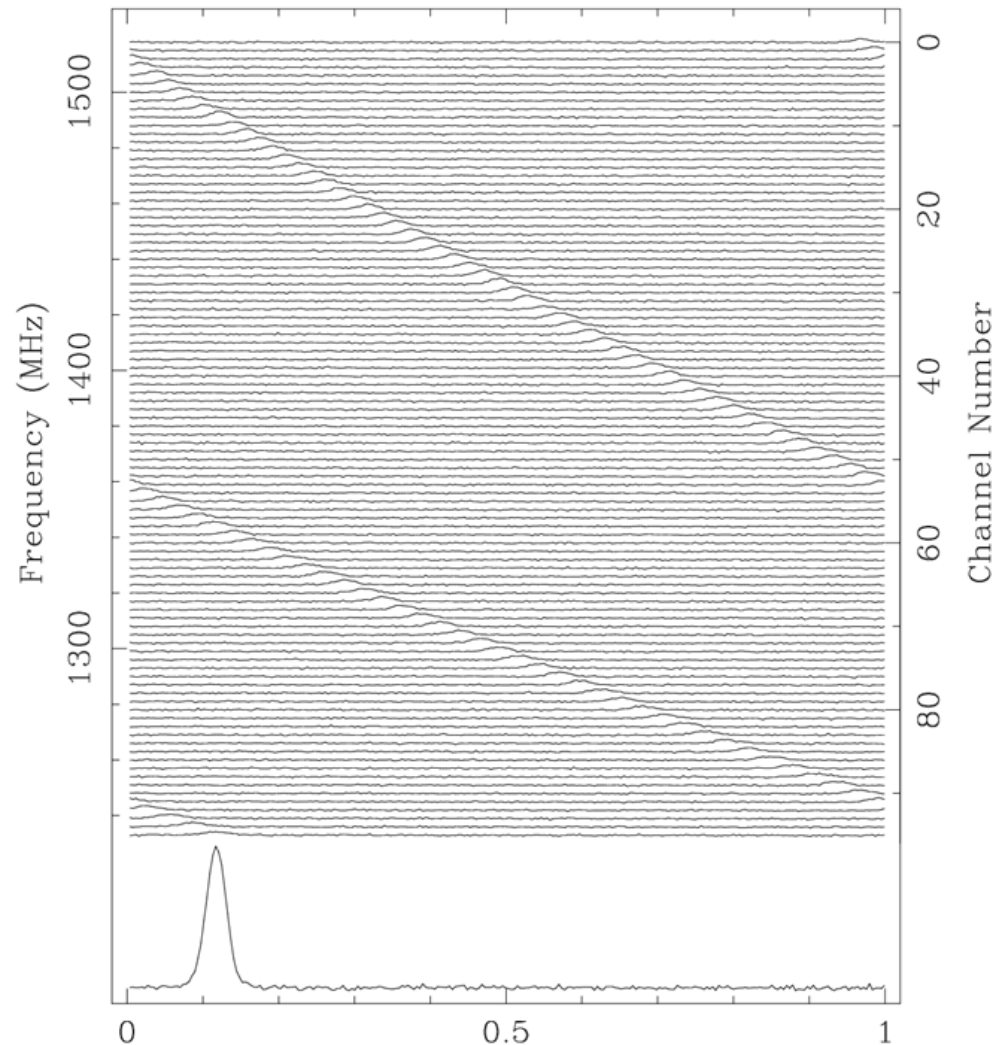
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Mid Pulsar Search

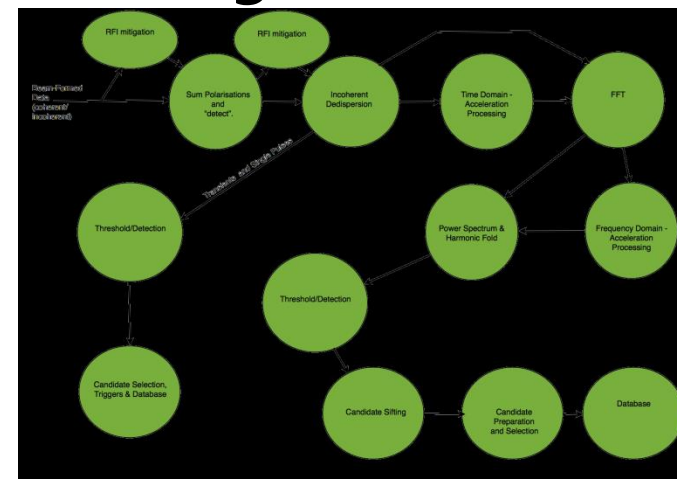
Pulsar Search Overview

- Search for pulsar signals:
 - Unknown periodicity
 - Signal deeply buried in noise
- De-dispersion:
 - Radio signals interact with free electrons, causes frequency dependent delay
 - Time signal searched over different dispersion measures to find candidate pulsars
- Acceleration search:
 - Particularly want to find pulsars in relativistic orbits around neutron stars or black holes
 - Period changes so guess acceleration



Pulsar Search Challenges

- Search for weak signals deeply buried in noise (ultimate needle-in-the-haystack)
- Simultaneously search 2222 beams
- Each beam has 4000-15000 channels, sampled every $50\mu\text{s}$ (700Gsample/s total) for 7min
- Three-dimensional search space amongst noise:
 - Unknown pulse period
 - Unknown dispersion
 - Unknown source acceleration



SKA₁ Key Computing-Imaging



3 SKA.TEL.DSHA - Dish Array

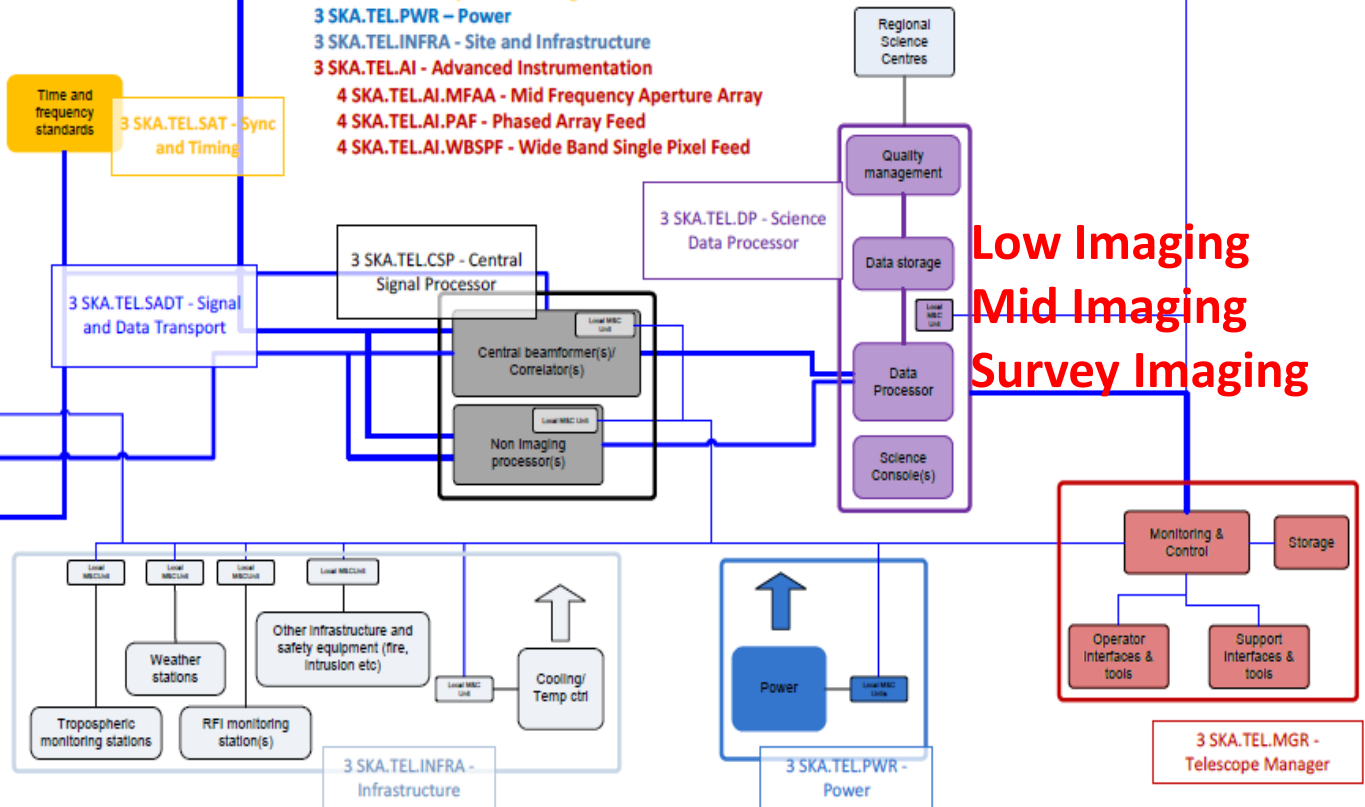
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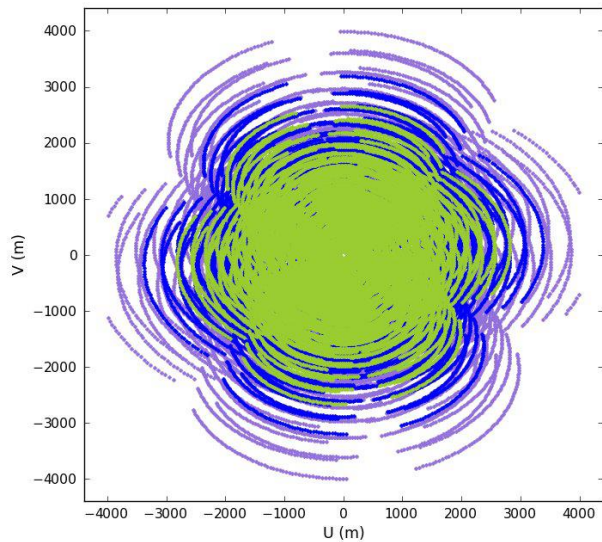


Low Imaging
Mid Imaging
Survey Imaging

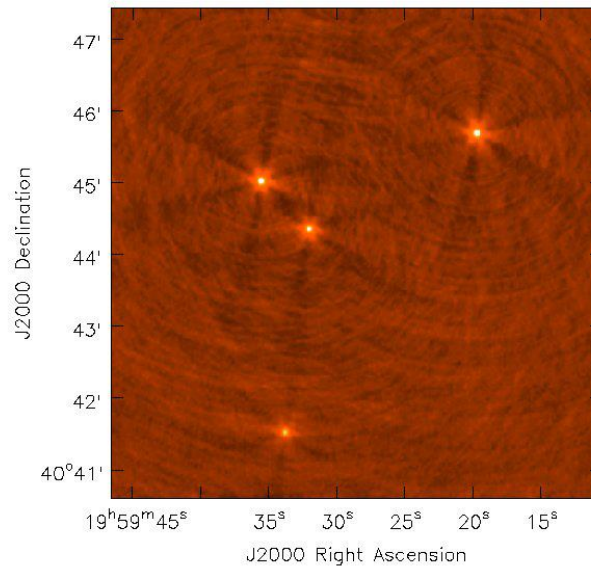
Imaging Overview



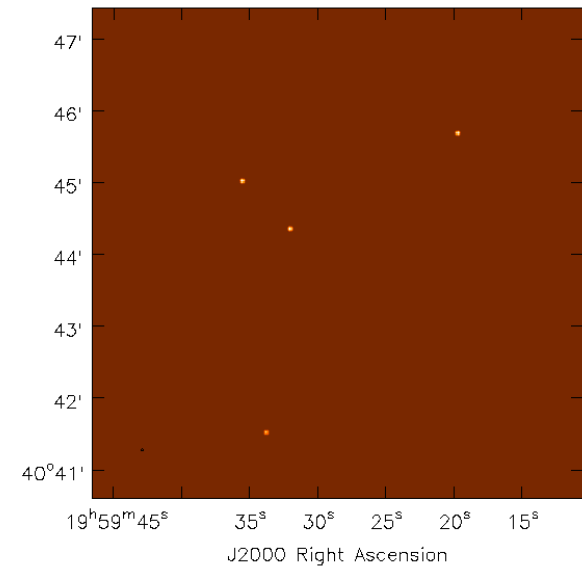
Visibilities from VLA correlator with 27 antennas over four hour observation



Dirty image after gridding visibilities and inverse FFT



Sought sky image after deconvolution



Imaging Challenges



- Gridding and (inverse) FFT are enormous compute challenges
- Using up to 60kx60k grid and including each of up to 2^{18} channels in the imaging pipeline requires supercomputing performance:
 - SDP estimated to require up to peak 100 PetaFLOP for Low, 350 PetaFLOP for Mid, 125 PetaFLOP for Survey (Cray Titan is 17.59 PetaFLOP), within reach by 2018 but everything must be done in real time 24-7
 - Power requirements determine ultimate costs, SDP computing need to fit within 5-10 MW
- SKA-2 will increase computing demands by approximately 100 fold
- Independent “compute island” design

SKA₁ Key Computing



Survey Beamformer

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Time and frequency standards and Timing

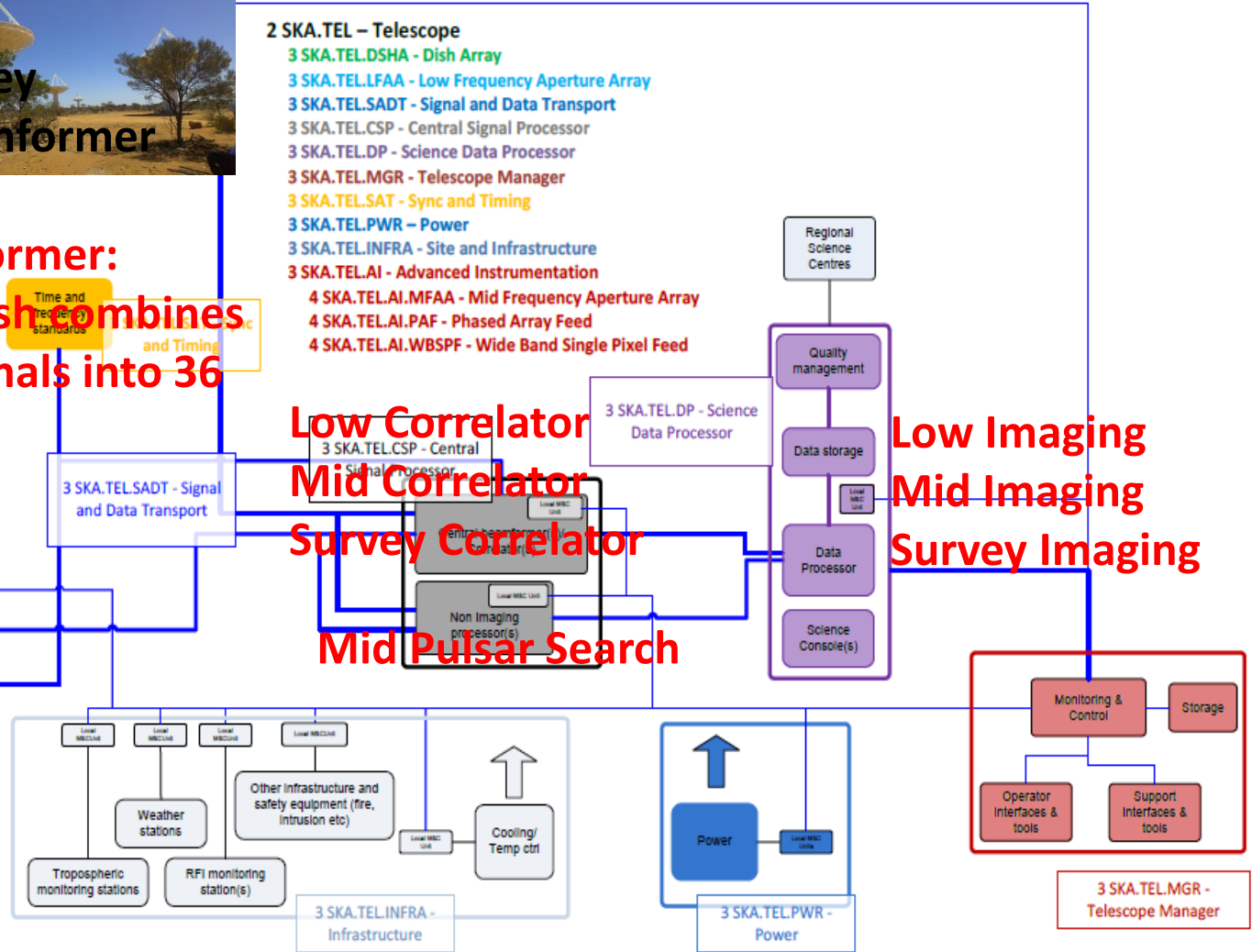
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Low Correlator
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Mid Pulsar Search

Low Imaging
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Survey Imaging

3 SKA.TEL.MGR - Telescope Manager

Questions?

Thank you

NZ SKA Work Packages

Work Packages	Lead	NZ Commitment
Correlators: - Survey - Mid & PowerMX - Low	NZ (Ensor) NRC (Carlson) Oxford (Zarb-Adami)	5.14 FTE/year 0.81FTE/year 0.45FTE/year
Imaging Pipeline	Southampton (Scaife)	1.1FTE/year
Pulsar Search	Manchester (Stappers)	0.70FTE/year
CSP & SDP Modeling	NZ (Wilson)	1.47FTE/year
CSP Prototyping	NZ (Ensor)	0.38FTE/year
SDP Computing - Hardware & Storage - Middleware - Common Software	Astron (Broekema) NZ (Bancroft) NZ (Christie)	0.64FTE/year 0.40FTE/year 0.61FTE/year
Software Development Environment	NZ (Erdody)	0.38FTE/year
Science Analysis Pipeline Requirements	NZ (Johnston-Hollitt)	2.50FTE/year