# New Zealand path to gamma-ray burst follow-up and transient observations Stuart Weston

Institute for Radio Astronomy & Space Research Auckland University of Technology

sweston@aut.ac.nz

### **SKANZ 2012**



Credit: Red Aurora - Alex Cherney

### Abstract

Due to its fast slewing rates, the Warkworth 12m radio telescope can be on source in a short period after a gamma-ray burst (GRB) alert and conduct follow-up radio observations of GRB and other transient events.

We report on the work that is currently being undertaken to integrate the Field System with the GRB Coordination Network (GCN) alert system.

The aim is to undertake both single-dish (Warkworth) and interferometric (with Australian antennas) observations of these events.

In conjunction with this project work is taking place to utilize real-time autocorrelating spectrometer with the aim to handle dedispersion when observing pulsars and searching for transient events.

AUT COMPUTING + MATHEMATICAL SCIENCES

### Acknowledgments

Keith Bannister CSIRO Astronomy and Space Science

Steven Tingay Curtin Institute of Radio Astronomy

Richard Dodson, International Centre for Radio Astronomy Research University of Western Australia

### GCN: The Gamma-ray Coordinates Network

Locations of GRBs and other Transients (the Notices) detected by spacecraft (most in real-time while the burst is still bursting and others that are delayed due to telemetry down-link delays).

An alerts is transmitted to the receiving sites via several methods, such as E-mail, Phone/Modem, Pager and Internet Sockets.



http://gcn.gsfc.nasa.gov/

### Example of alert received through GCN socket\_demo

 PKT INFO:
 Received: LT Thu Dec 8 01:22:23 2011

 Type= 112
 SN= 1
 FERMI-GBM Gnd Position

 Hop\_cnt= 0
 PKT\_SOD= 0.00 [sec] delta=44543.00 [sec]

 RECORD\_NUM=
 56

 TRIGGER\_NUM=
 344953038

 TJD=
 15902

 SOD=
 44236.20 [sec] delta=306.80 [sec]

 RA=
 166.8100 [deg] (J2000)

 Dec=
 -18.9700 [deg] (J2000)

 ERROR=
 10.58 [deg radius, statistical only]

 DATA\_SIGNIF=
 5.20 [sigma]

 DATA\_INTERVAL=
 1.024 [sec]

 PHI=
 51.00 [deg]

 THETA=
 28.00 [deg]

 E\_RANGE=
 44.032 - 279.965 [keV]

 LOC\_ALGORITHM=
 4143 (Gnd S/W Version number)

 This is a ground-generated Notice -- not flight-generated.

I have put socket\_demo under "/usr2/st/gcn" on the Field System. As this will be a local station procedure. It has been modified to also communicate with the DBBC and Mark5B.

### **Current System Configuration**



## AUT 12m Antenna

Antenna type	: Fully steerable dual shaped Cassegrain
Main dish Diam	: 12.1 m
Slewing rates	: (5 deg/s in azimuth and 1.25 deg/s in elevation)
Acceleration	: (1.3 deg/s/s)
Secondary Refl Diam	: 1.8m
Focal Length	: 4.538 m
Surface precision	: 0.35 mm (RMS)
Wavelength	: S-Band and X-Band
Mount	: alt-azimuth

#### AUT COMPUTING + MATHEMATICAL SCIENCES

### **Field System log for Alert**

2012.038.09:19:59.37;Log Opened: Mark IV Field System Version 9.10.4 2012.038.09:19:59.37;location,WARKWRTH,185.34,-36.43,132.0 2012.038.09:19:59.37;horizon1,0.,10.,360. 2012.038.09:19:59.37;antenna,12.0,300.0,60.0,-180.0,360.0,8.0,88.0,azel 2012.038.09:19:59.37;equip,none,mk5b bs,none,none,500.10,3,a/d,101,60,20,none,40,0,in,8bit,cdp,3,return 2012.038.09:19:59.37;time,-5.459,1.011,rate 2012.038.09:19:59.37;flagr,200 2012.038.09:19:59.37;source=134595401,000000.00,-350000.00,2000.0,

2012.038.09:19:59.40#antcn#Commanding to a new source 2012.038.09:19:59.41#antcn#/case1/Position Requested. RA: 0.154267, DEC: -34.934311

2012.038.09:19:59.41#antcn#/pmodl/azcor = -0.514445

2012.038.09:19:59.41#antcn#/pmodl/elcor = 1.183612

2012.038.09:19:59.41#antcn#/pmodl/AzEl:239.505301 20.071473 2012.038.09:19:59.41#antcn#/casel/Pointing Model Corrections -0.514445 1.183612

2012.038.09:19:59.41#antcn#/case1/Corrected RA & DEC 1.083632 -35.965338

2012.038.09:19:59.41#antcn#Turning power ON. 2012.038.09:20:00.10#flagr#flagr/antenna,new-source 2012.038.09:21:17.21#flagr#flagr/antenna,acquired 2012.038.09:25:07.56;log=station

### 1min 18 secs to be on source and tracking

### GPU's the next step in compute power!



### **Future System Configuration**



### Real-time autocorrelating spectrometer

Stream-based approach using hardware accelerators for radio astronomy applications, Mahmoud Mahmoud, AUT

Future work is to link this with the Field System

Also to accept real-time streaming of data either from the DBBC or Mk5C

We are also working with Richard Dodson and another GPU Autocorrelator (http://sourceforge.net/projects/g-auto/)

This has been tested real-time over the network from Warkworth (via KAREN) to Perth (via AARNET).

Future work is to link this with the Field System

Comparison between a multi-core CPU (IBM blade) server and one GPU.

IBM Blade -m5spec 16MHz Band with 32767 channels, 12 mins GPU – autocorrelator same data file but with 524290 channels, 9.8 secs

#### 🕅 Xming :10.0

Applications Places System 🤗



### Conclusions

>Good things can be done with a small dish

>We can be on source within 2 minutes

>GPU's provide a significant leap forward for real-time processing

### References

BANNISTER, K. 2011, 'Results of automatic, high time resolution GRB follow-up with the Parkes 12 m', ATUC Science Day - Time Domain Astronomy, ATNF

DODSON, R., PHILLIPS, C., FERRIS, D. 2009, 'A GPU based solution for distributed FX correlation', 8th eVLBI Workshop Express 09, Italy, 82, pp. 37

Southern Cross

The CoalSack Dark Nebula

Large Magellanic Cloud

Alpha Centauri

Antares

Small Magellanic Cloud

47 Tucanae Globular Cluster

Achernar

Credit: Red Aurora - Alex Cherney

# Thank You



#### AUT COMPUTING + MATHEMATICAL SCIENCES