Precision Effects of Calibration in the Imaging Pipeline

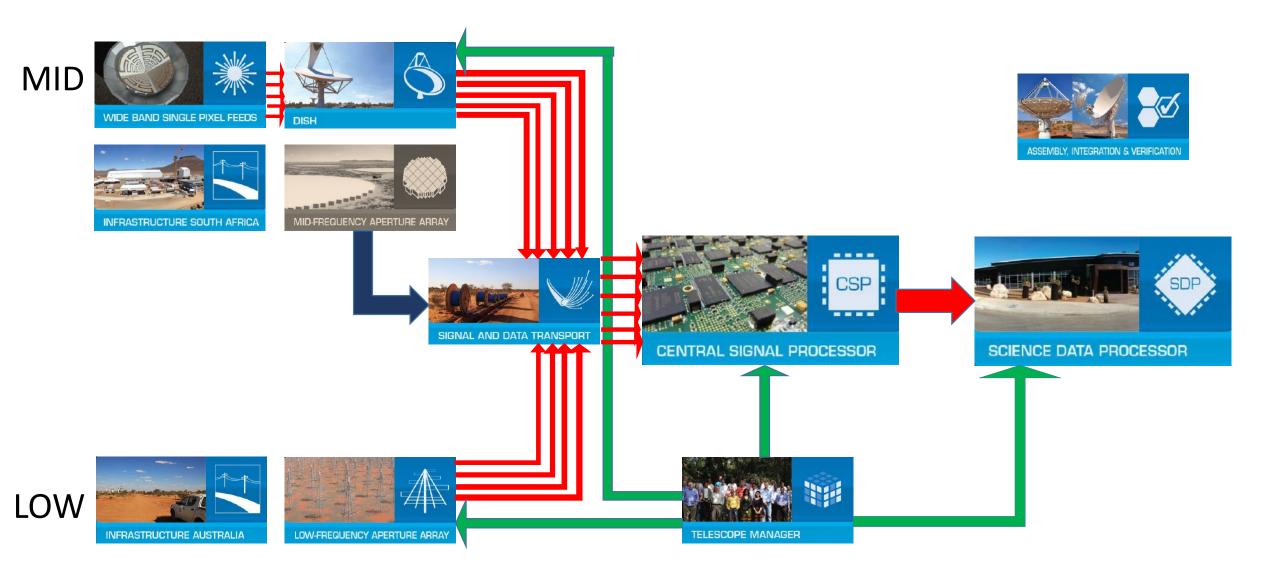
14 February 2019

Anthony Griffin, AUT



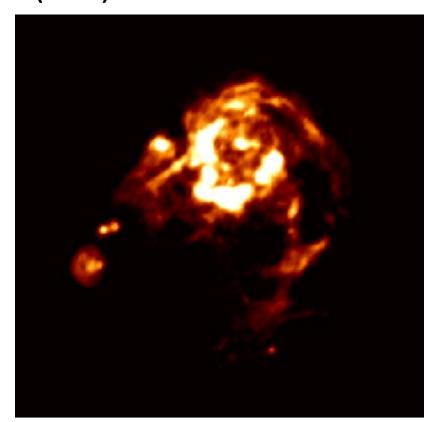


SKA Phase One Data Flow and Consortia Teams

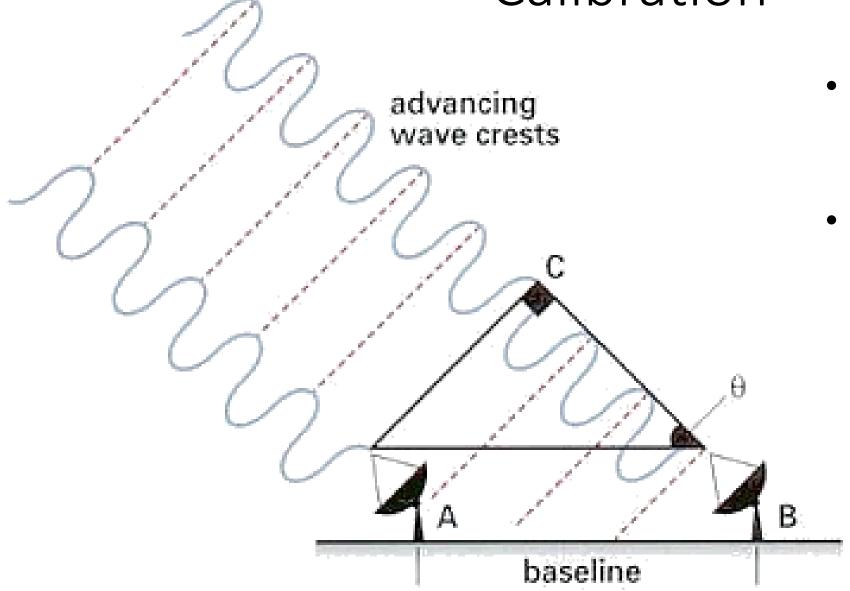


Imaging Pipeline

- Major function of the Science Data Processor (SDP)
- Takes the output of the Channel Signal Processor (CSP)
 - Visibilities (uv-plane)
 - (Measurements of the sky in the Fourier domain)
- Processes the Visibilities
- Produces an image of a region of the sky

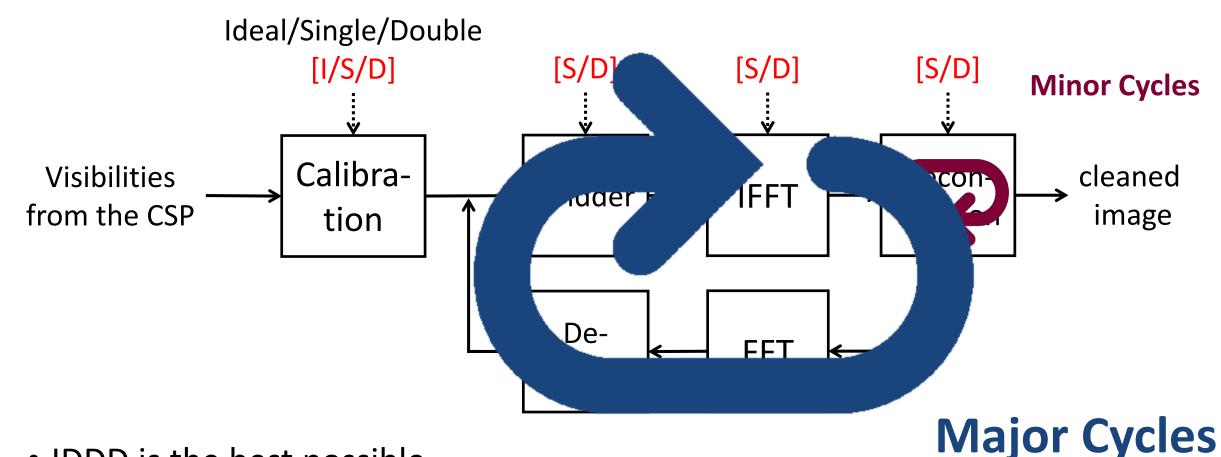


Calibration



- Many things need to be calibrated in a radio inteferometer.
- So far we are just looking at amplitude calibration.

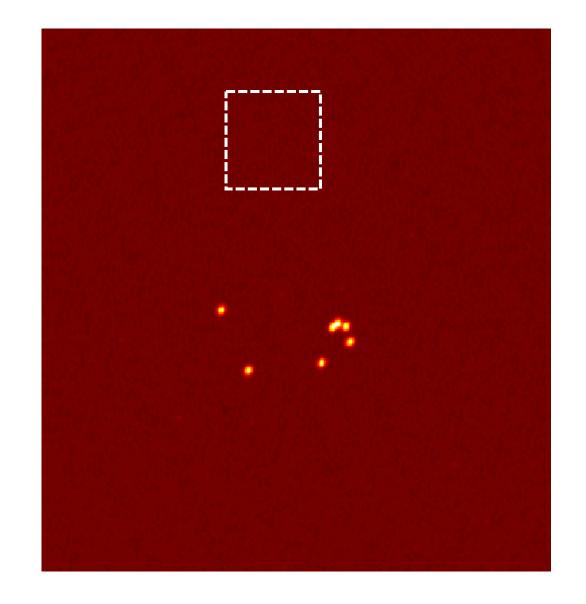
Precision Options in the Model



- IDDD is the best possible
- For example, DSDD vs DDDD shows the effect of gridding in single precision

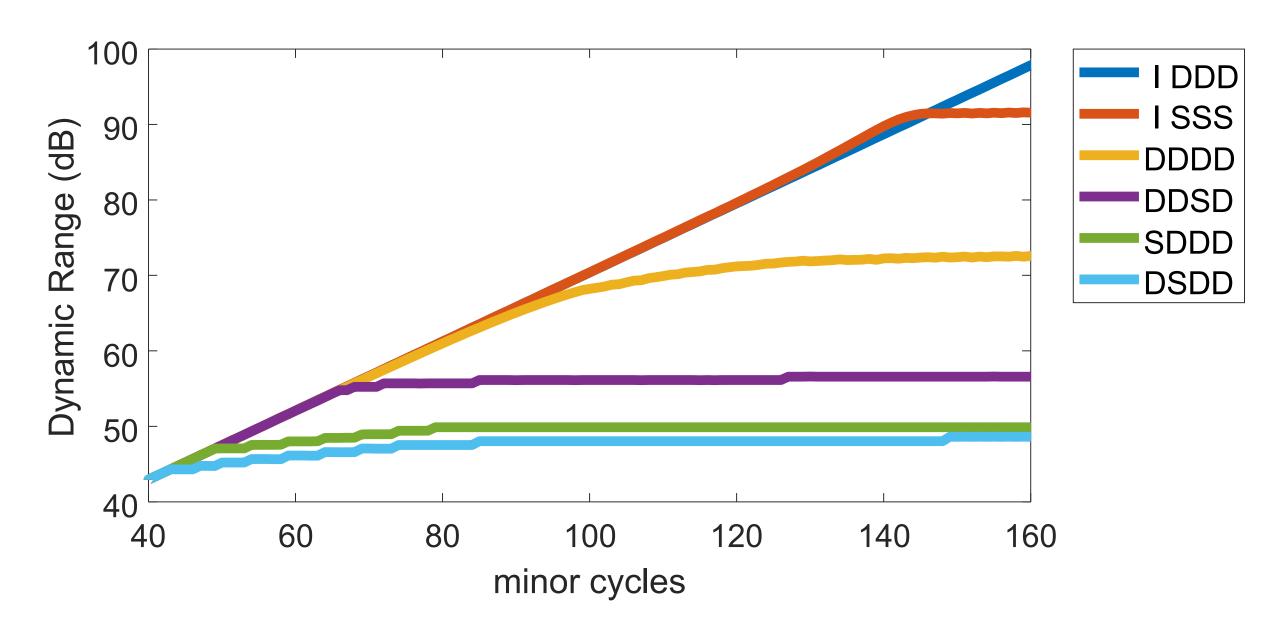
Dynamic Range

$$Dynamic Range = \frac{Peak Value}{Energy in a Patch}$$



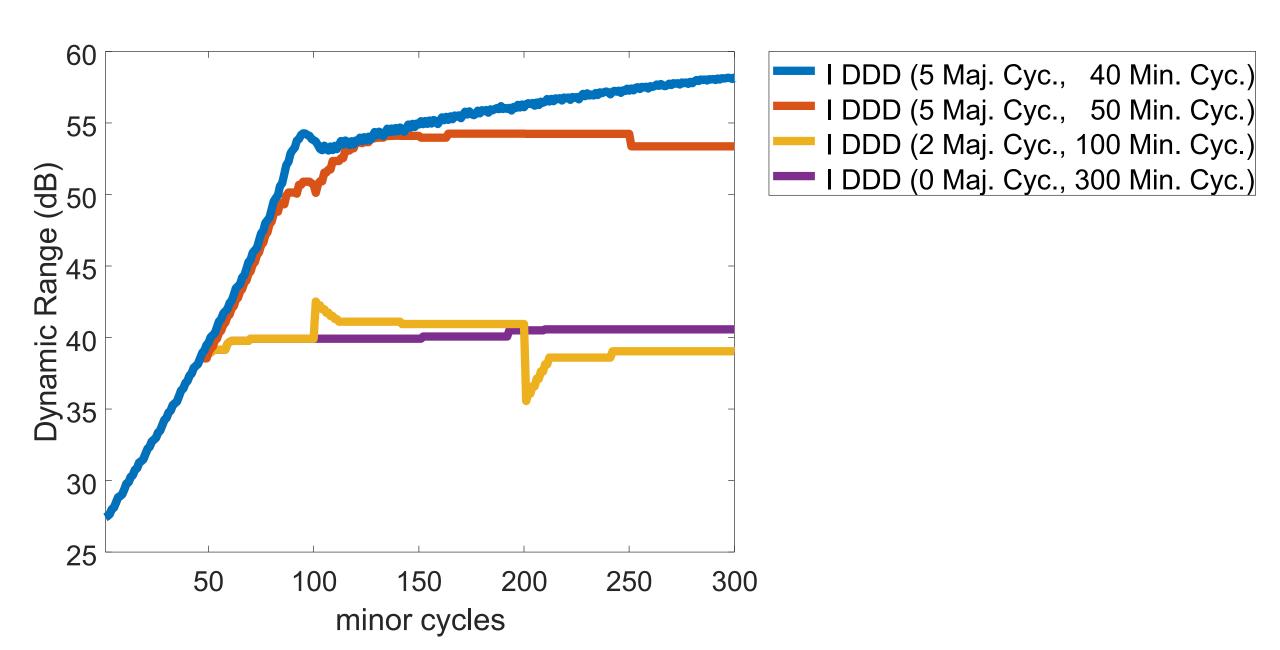
Simulation Parameters

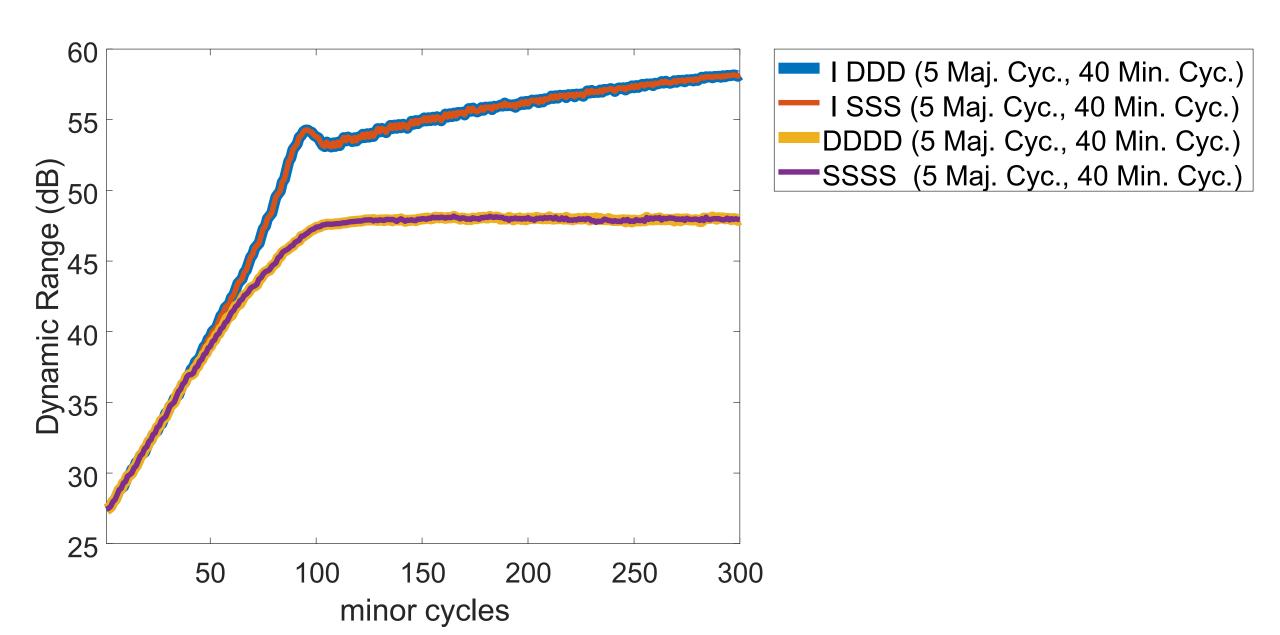
- SKA-LOW -> 512 receivers 130,816 baselines
- 1 second observation time
- 1 source at the phase centre
- 1 fine channel
- 200 cycles of Högbom cleaning, with a gain of 0.1
- Amplitude calibration with an SVD solver



Simulation Parameters

- SKA-LOW -> 512 receivers 130,816 baselines
- 1 second observation time
- 2 sources
 - 1 close to the phase centre, the other further away, neither on a pixel
- 1 fine channel
- 5 major cycles and 40 minor cycles of Högbom cleaning
 - a gain of 0.1
- Amplitude calibration with an SVD solver





Conclusions and Future Work

- Single precision on calibrated data does have some effect on Dynamic Range (DR) in the final image
- In the extreme case (point source at phase centre)
 - FFTs in single precision result in a DR ceiling of about 57 dB
 - Gridding in single precision results in a DR ceiling of about 48 dB
- No change in the two source test
- These are preliminary results!
- Longer observations
- More complicated skies
- More realistic Calibration algorithms
 - Self-cal, Stef-cal, etc