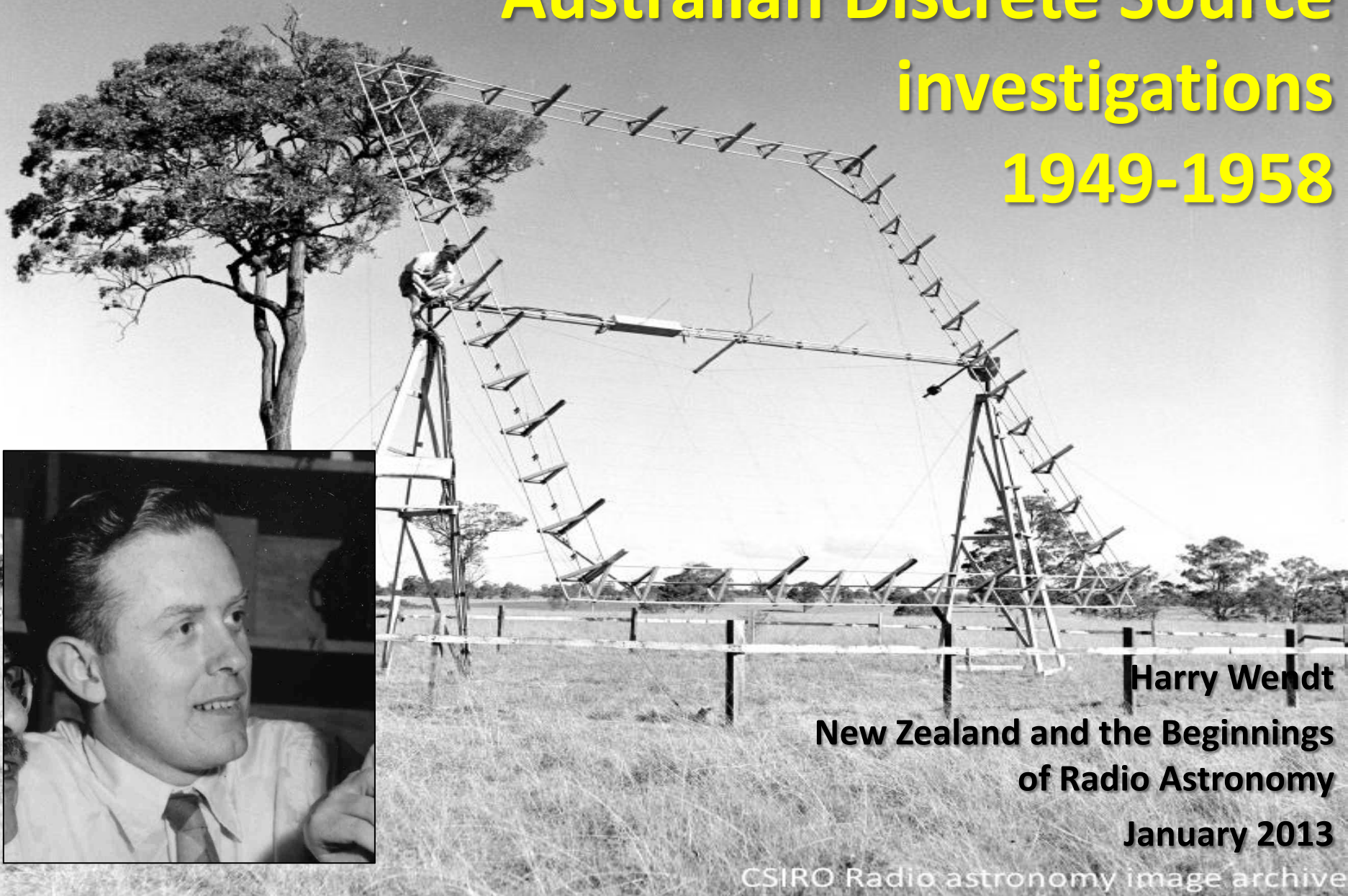
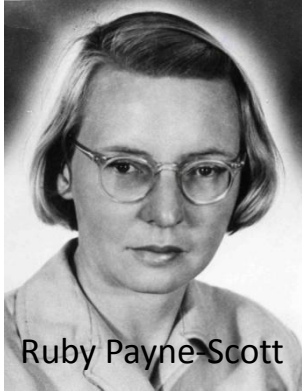


Bernard Mills and the other Australian Discrete Source investigations 1949-1958

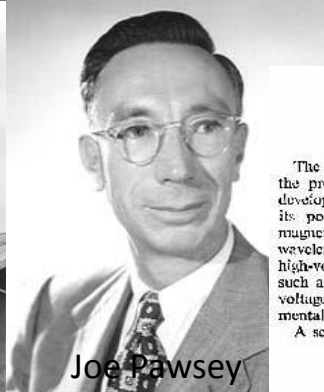


Harry Wendt
New Zealand and the Beginnings
of Radio Astronomy
January 2013

Background on Bernard Mills



Ruby Payne-Scott



Joe Pawsey



CSIRO Radio astronomy image archive

A MILLION-VOLT RESONANT-CAVITY X-RAY TUBE

621.386:1.027.89; 621.396.611.4 Paper No. 924 RADIO SECTION

B. Y. MILLS, B.Sc., M.E.

(Draft of a paper which will be published in Part III of the Proceedings.)

The idea of utilizing resonance in a high-frequency circuit for the production of high-voltage particles is not new, but the development of radar techniques during the war greatly increased its possibilities. The high power available from a pulsed magnetron, coupled with the small size of components at microwave-lengths, permits of the construction of an extremely compact high-voltage source. The paper describes the construction of such a device intended for operation as an experimental high-voltage X-ray tube. Some of the technical problems and fundamental limitations of the method are discussed.

A schematic of the apparatus is shown in Fig. 1. Briefly it

is injected from a point-focused gun having a peak current capability of about 400 mA. They are accelerated across the gap between the re-entrant cones by the r.f. field and strike a thin gold target of the transmission type. The X-ray beam is confined by a lead shield surrounding the cavity to a narrow cone in the direction of the electron beam. The peak voltage of the electron beam is monitored by coupling a small amount of the cavity power into a calibrated thermistor bridge.

A mean beam current of 70 μ A at a peak voltage of 1.1 MV has been obtained, whilst the highest recorded voltage is 1.25 MV with a very low beam current. An attractive feature

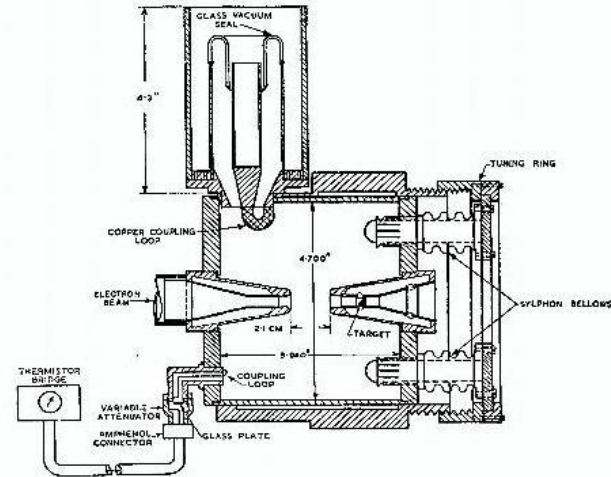


Fig. 1.—Assembly drawing of the cavity.

consists of a re-entrant copper cavity resonant in the region of 25 cm and fed by a 25-cm magnetron through a coaxial line and vacuum-tight choke joint. The cavity is continuously evacuated. The magnetron has an available power of about 600 kW with a pulse length of 5 microseconds and a pulse repetition frequency of 200 r/s. Because of difficulties associated with coupling a self-excited oscillator to a high-Q resonant circuit, it has only been possible to dissipate some 300 kW in the cavity. Electrons are

injected from a point-focused gun having a peak current capability of about 400 mA. They are accelerated across the gap between the re-entrant cones by the r.f. field and strike a thin gold target of the transmission type. The X-ray beam is confined by a lead shield surrounding the cavity to a narrow cone in the direction of the electron beam. The peak voltage of the electron beam is monitored by coupling a small amount of the cavity power into a calibrated thermistor bridge.

An important possibility, although one which has not been explored, is that the tube could be used as a stroboscope to radiograph machinery undergoing a cyclical motion. The

Mr. Mills is in the Division of Radiophysics, Commonwealth Scientific and Industrial Research Organisation, Australia.



1948 Eclipse

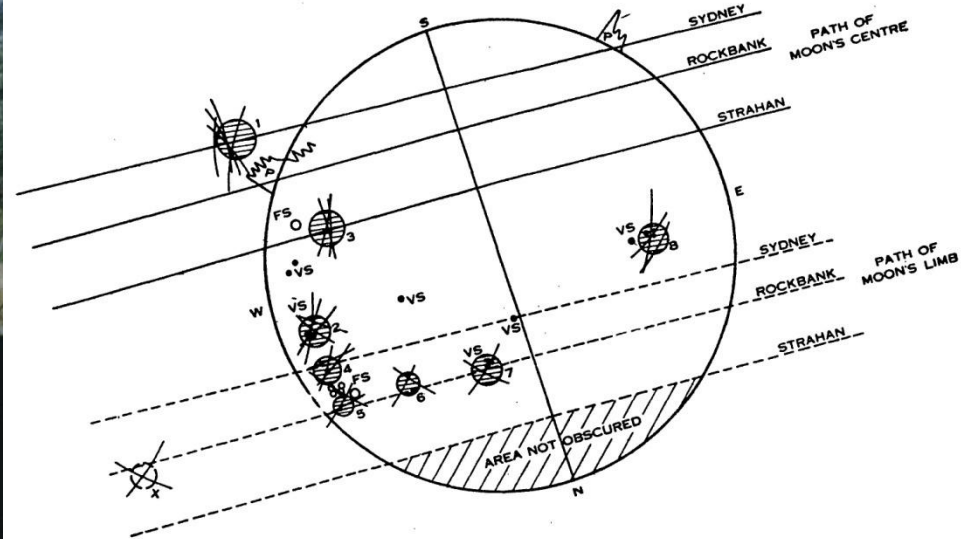
Christiansen, Yabsley & Mills



"Chris" Christiansen



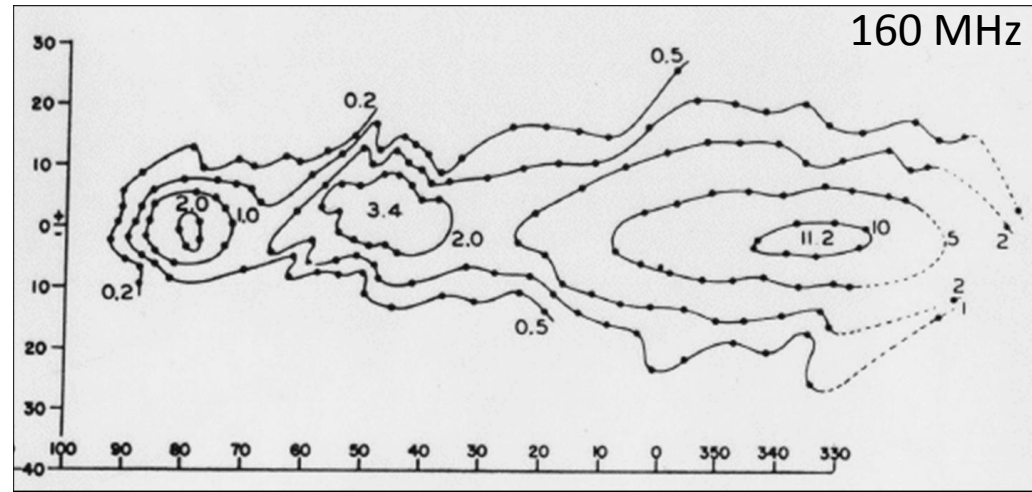
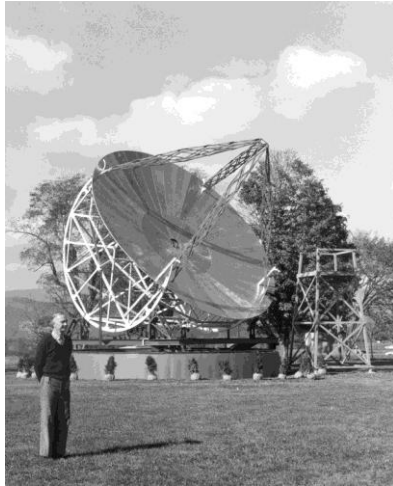
Don Yabsley



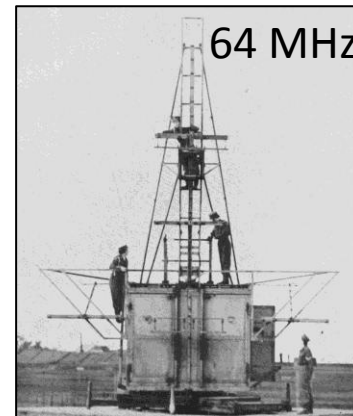
LIFE

Background on Discrete Radio Sources

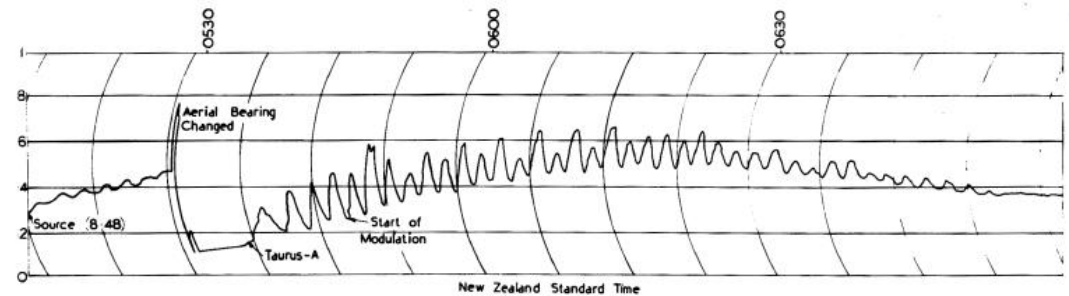
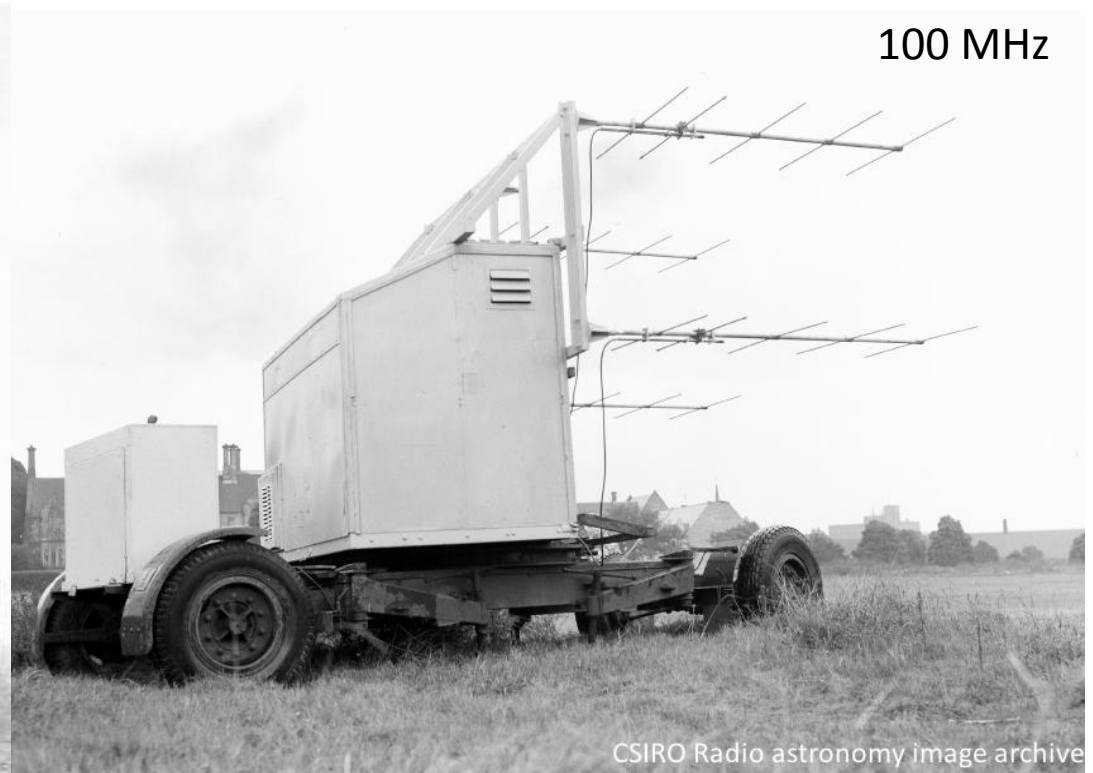
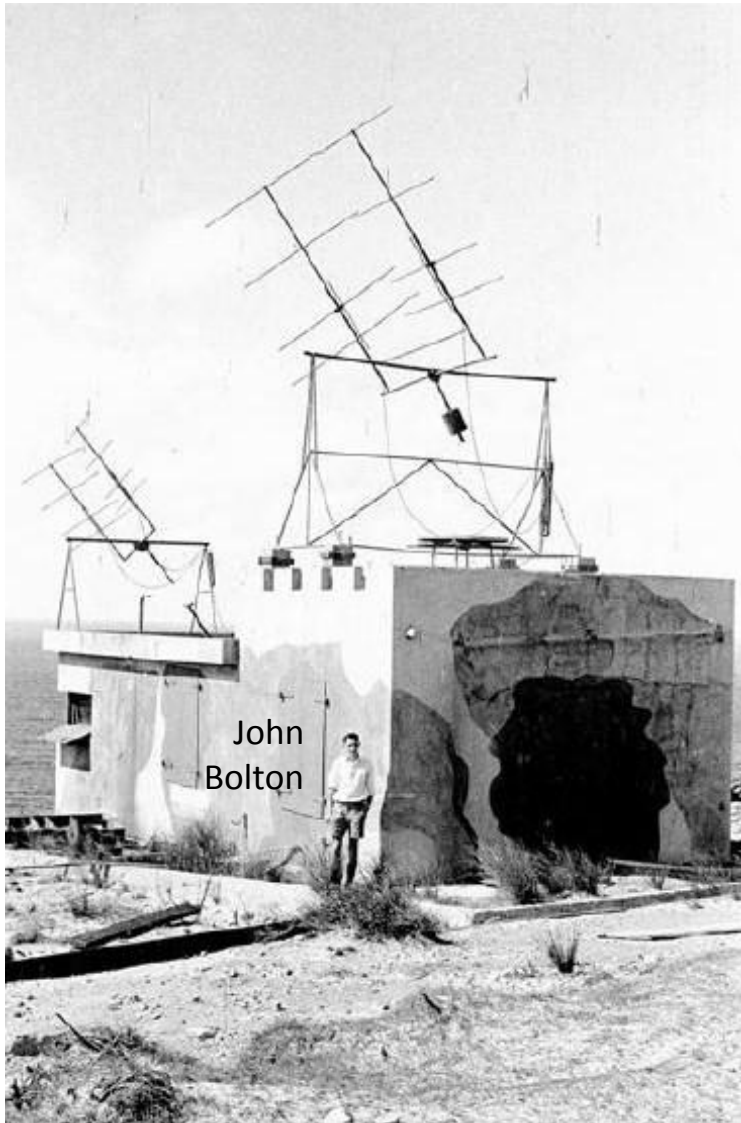
Reber 1944



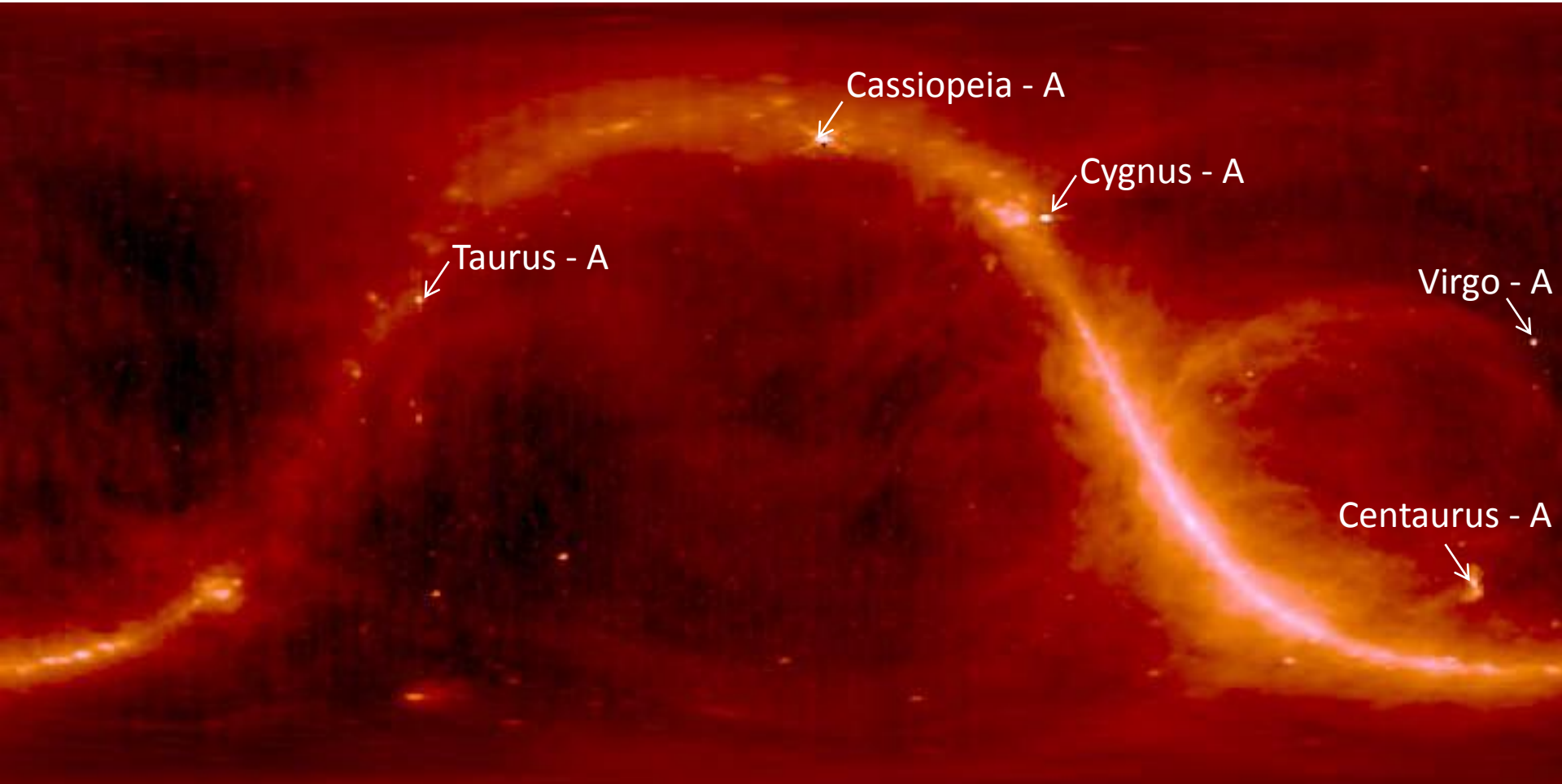
Hey, Parsons & Phillips 1946



Australian Discrete Sources Research



408 MHz Sky



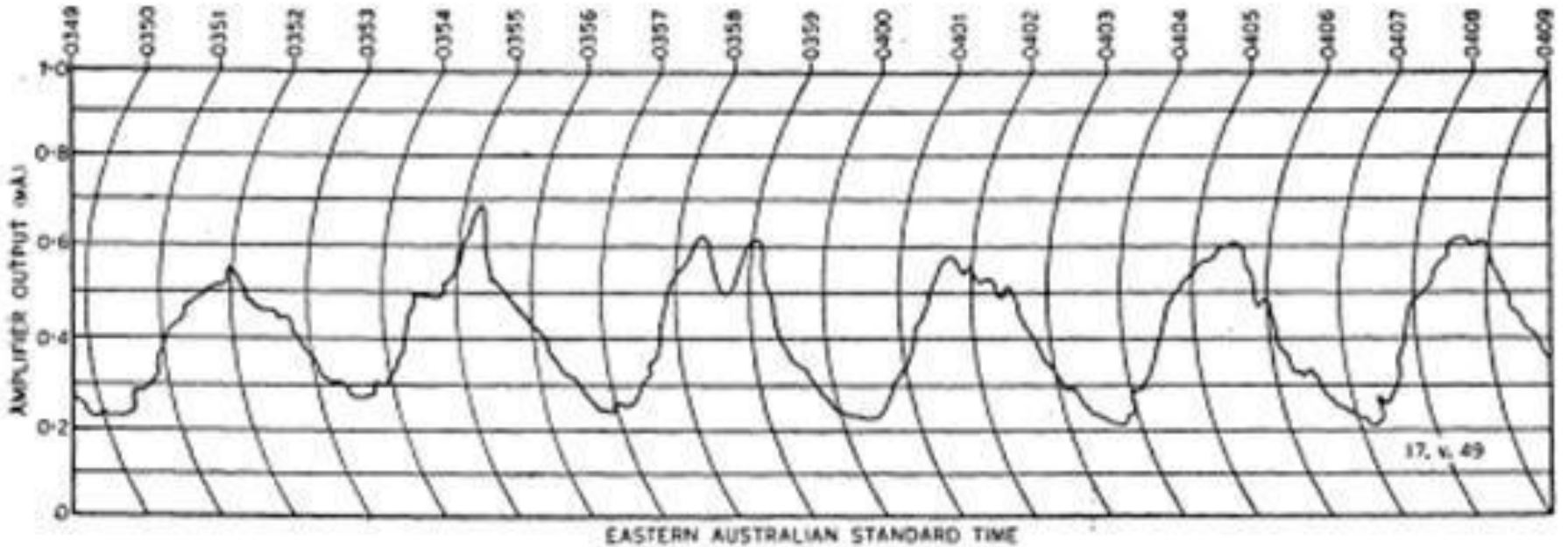
Potts Hill 1949

- 97 MHz Observations with Aidan Thomas
- Cygnus A position and tentative identification discounted
- Ionospheric scintillation confirmation



POTT'S HILL INTERFEROMETER SITE LOOKING NORTH-EAST

Cygnus-A 97 MHz



Field Station Locations

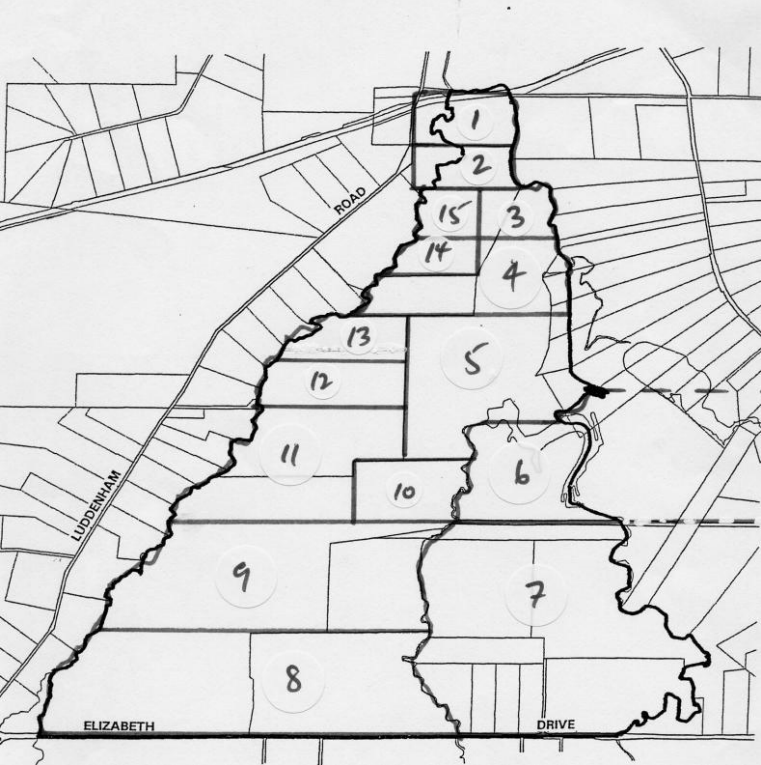
● **Badgerys
Creek**

● **Potts Hill**

Data SIO, NOAA, U.S. Navy, NGA, GEBCO

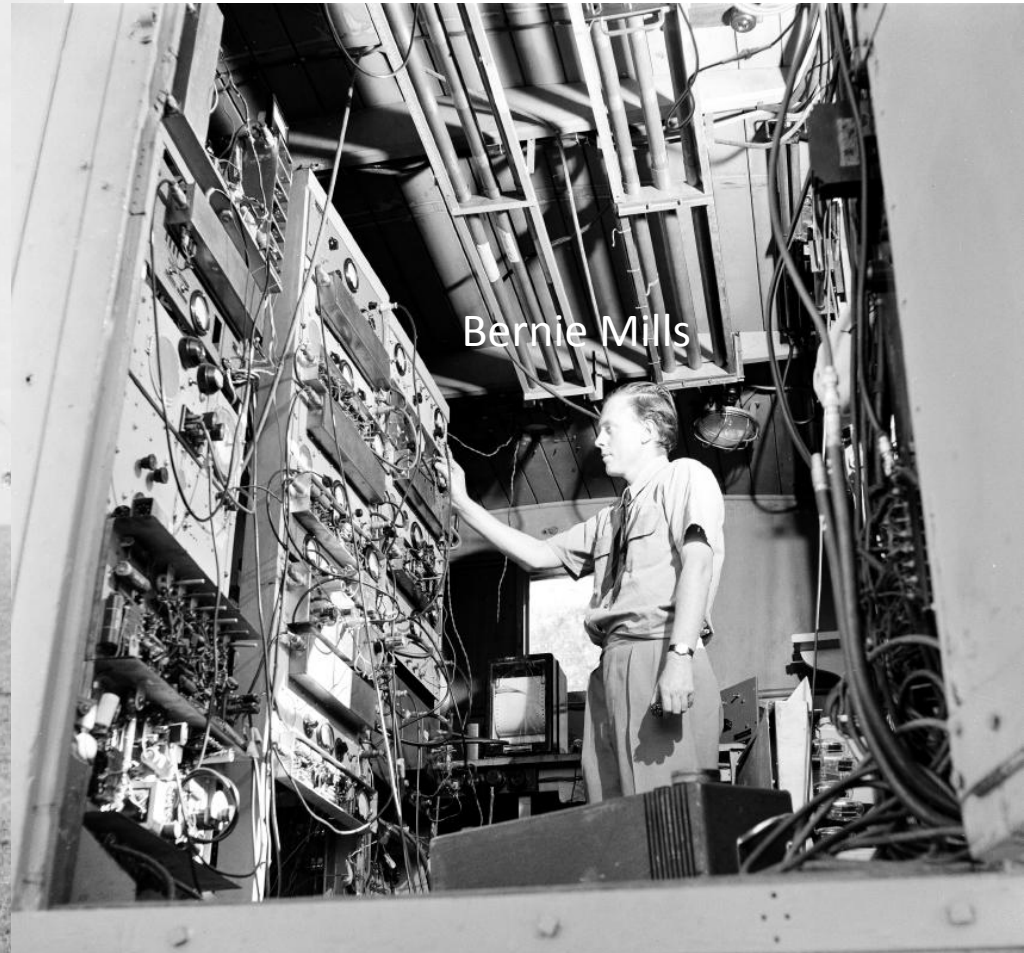
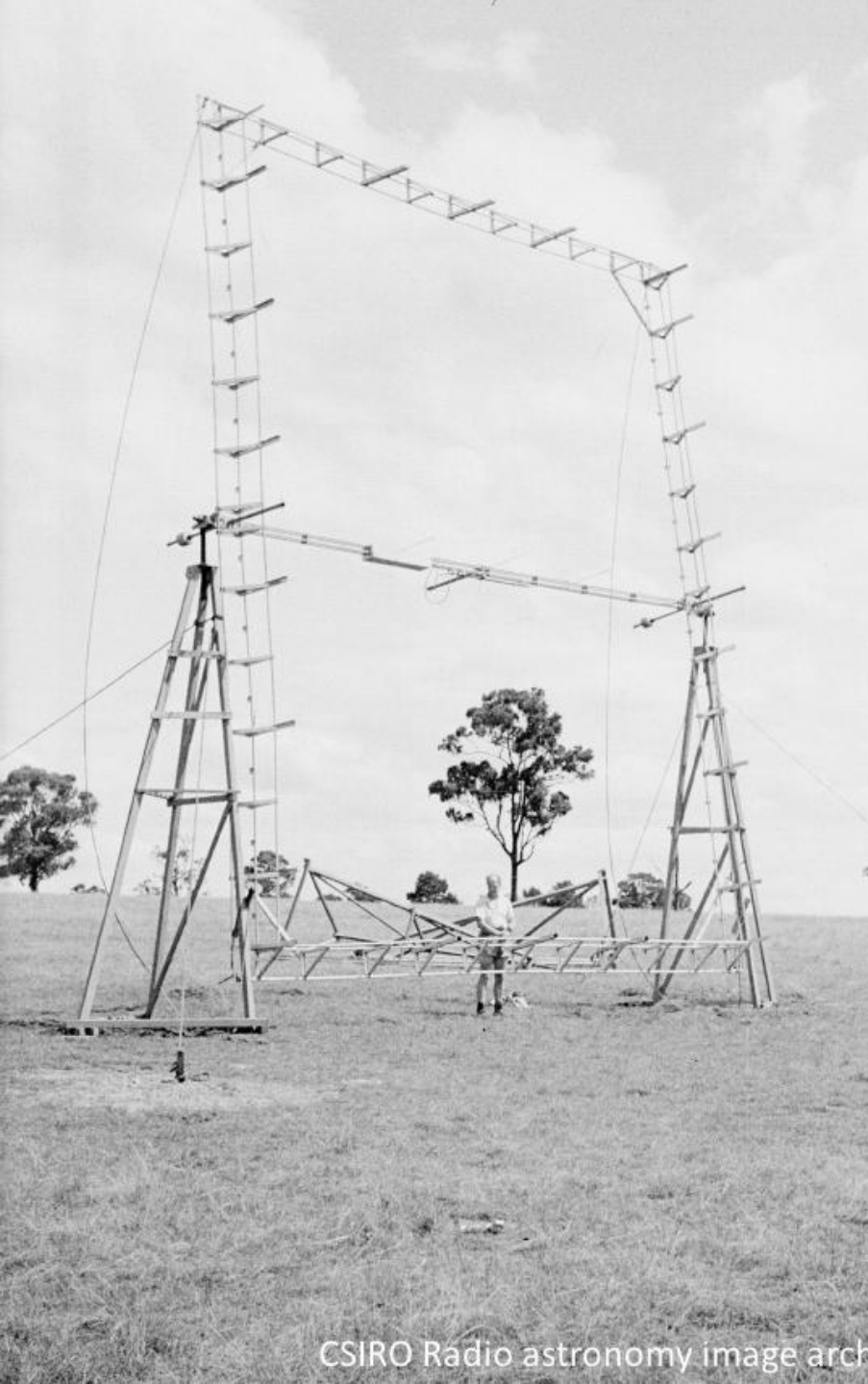
Image © 2012 Sinclair Knight Merz

Google earth



Badgerys Creek

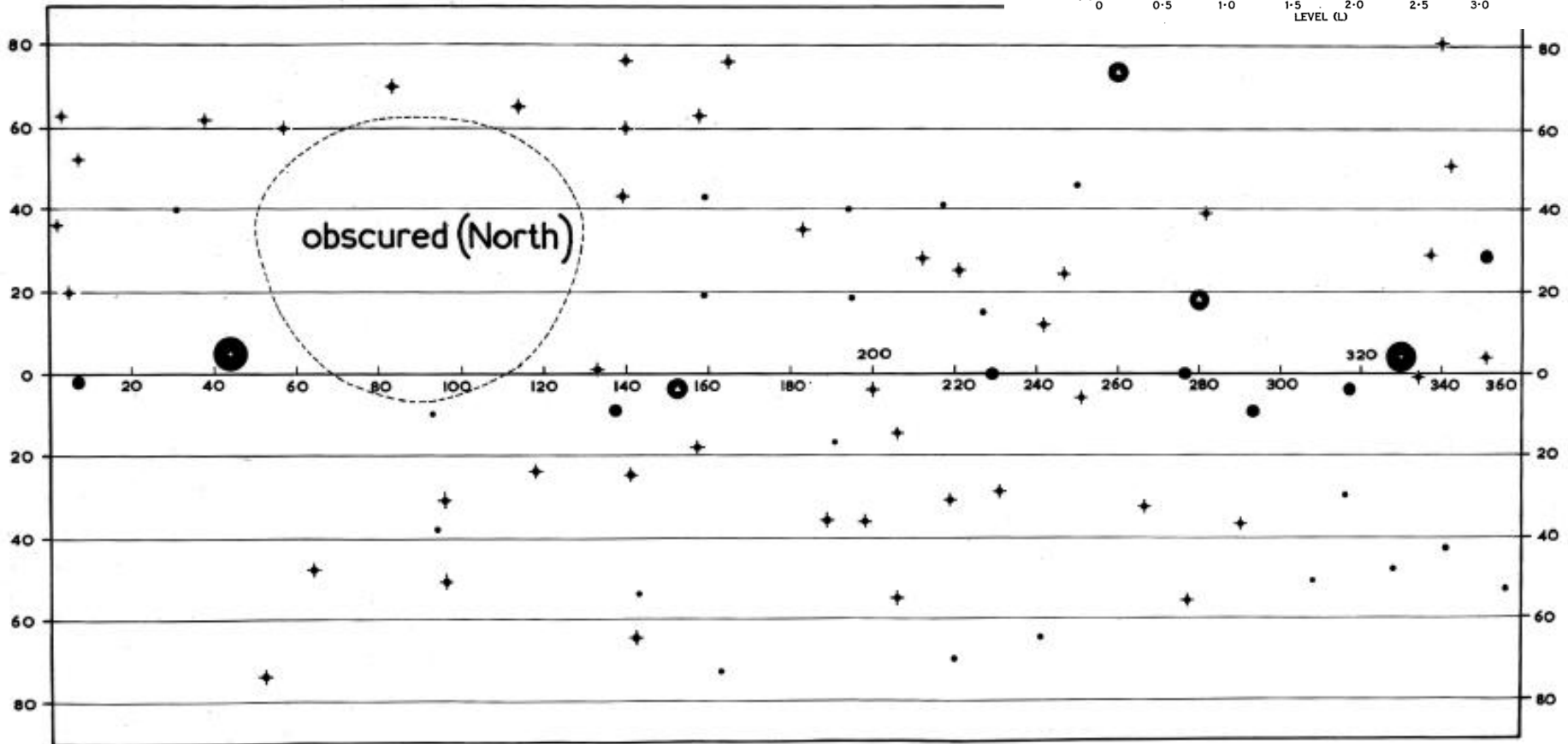
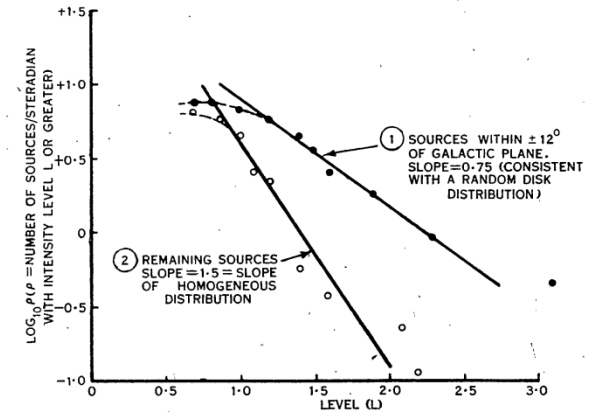
- 101 MHz Observations supported by Arthur Watkinson

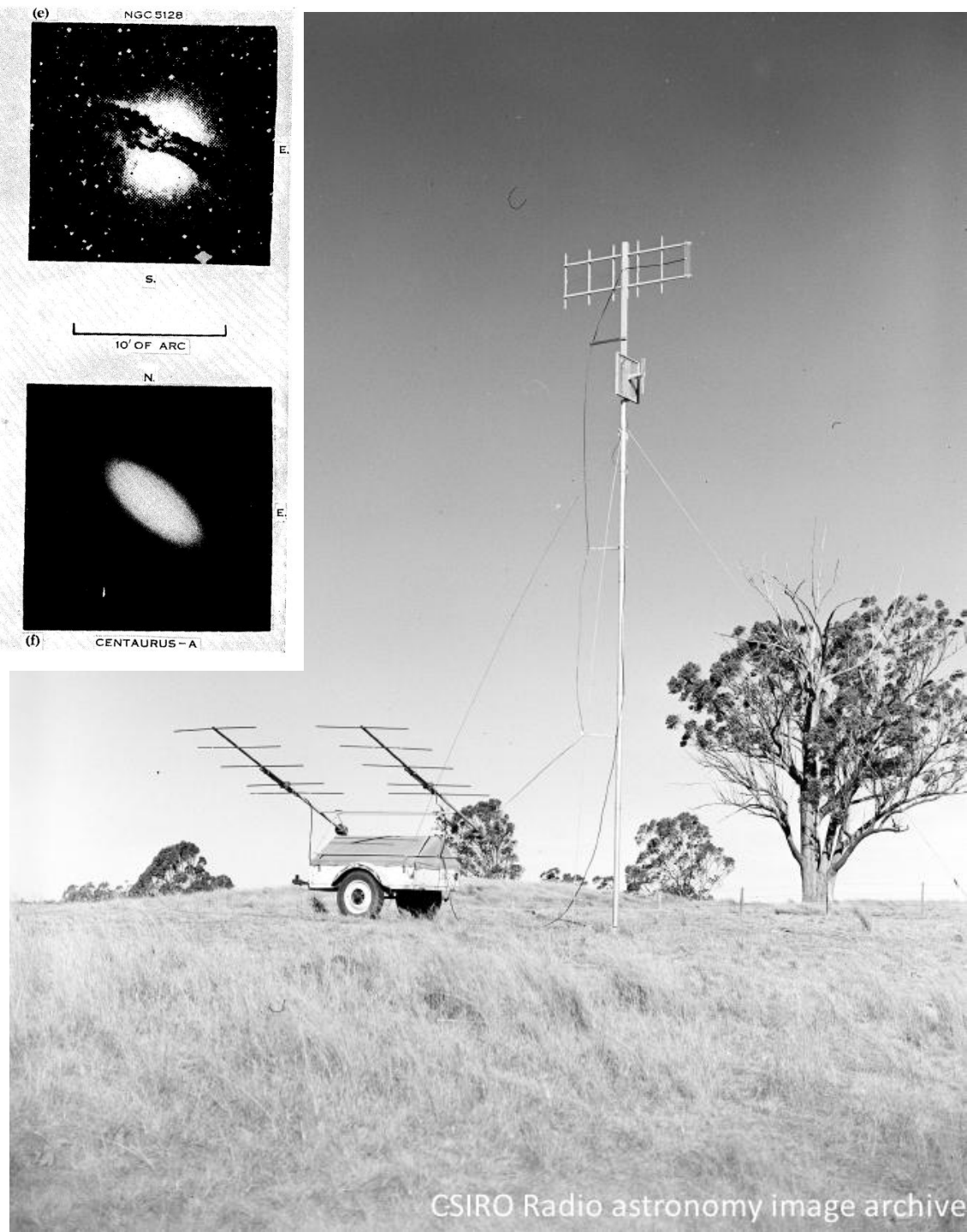
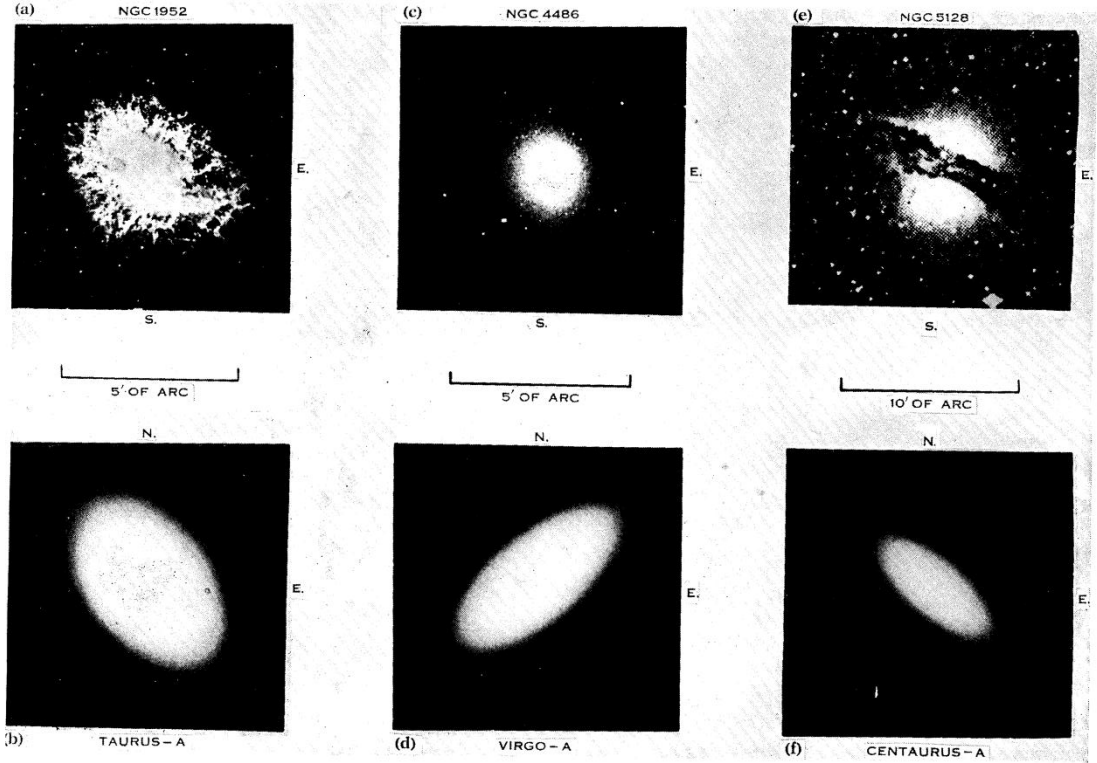


Bernie Mills

Discrete Source Survey 1952

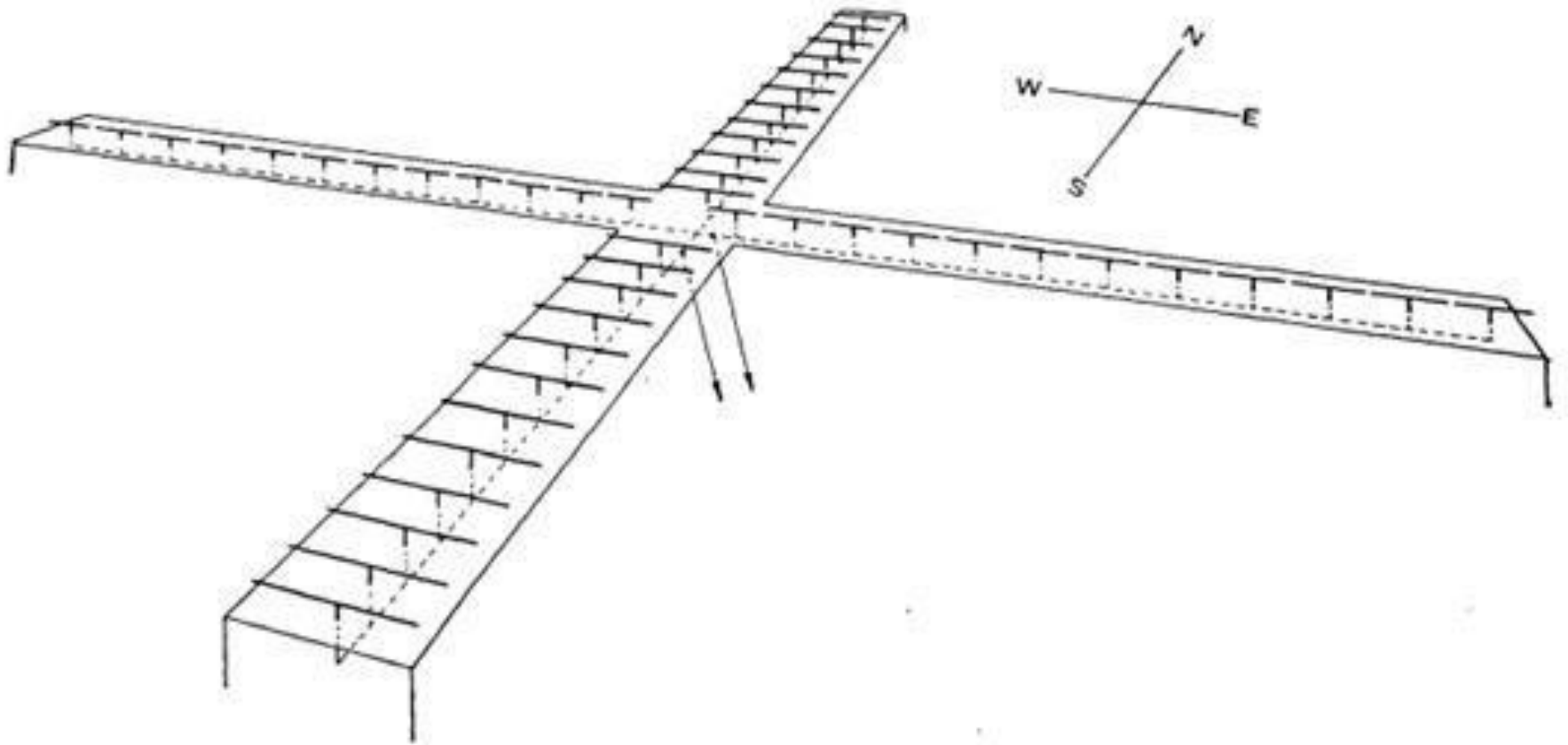
- Confirmed Bolton, Stanley & Slee Discrete Source Identifications
- 77 Discrete sources, two main population classes





- Radio-Link Interferometer
- Radio Brightness Distributions
- Preliminary results at URSI 1952

The Idea for a Cross

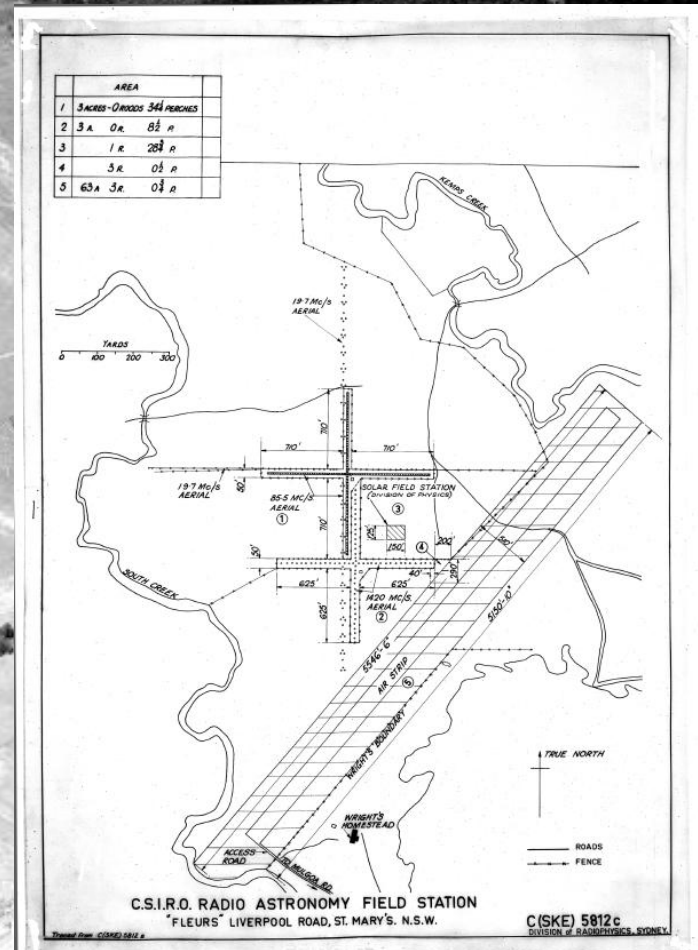
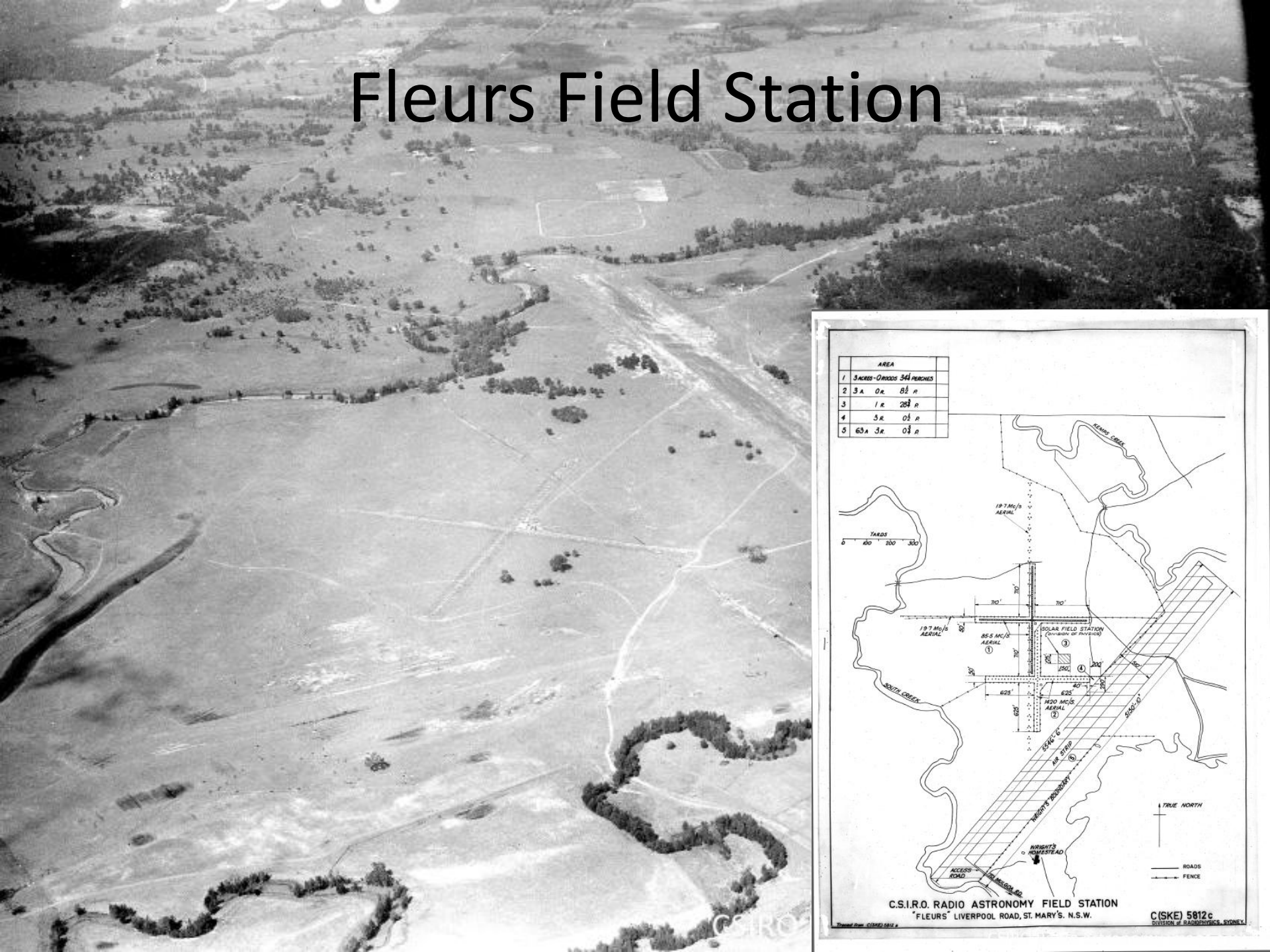


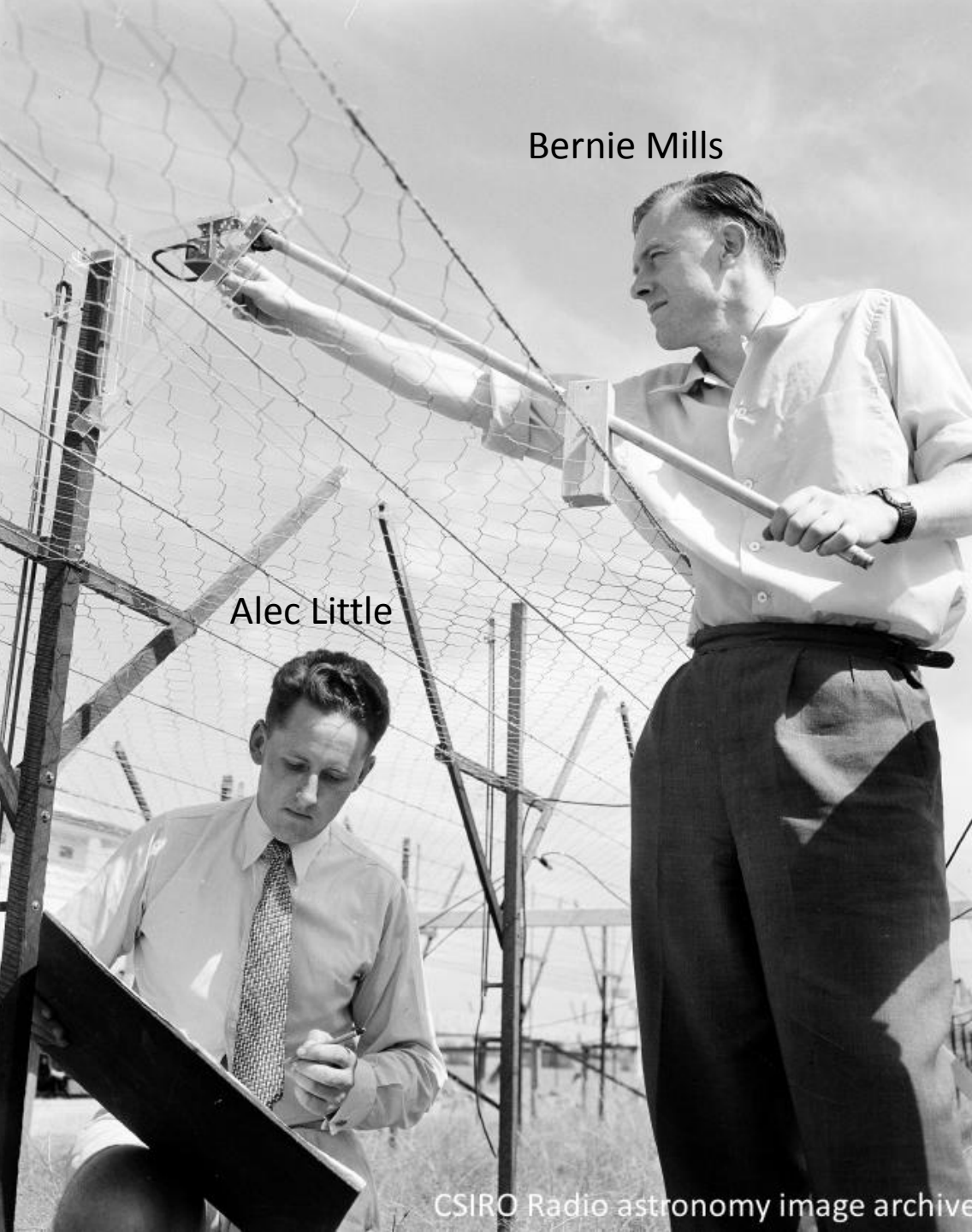
Prototype at Potts Hill 1953





Fleurs Field Station





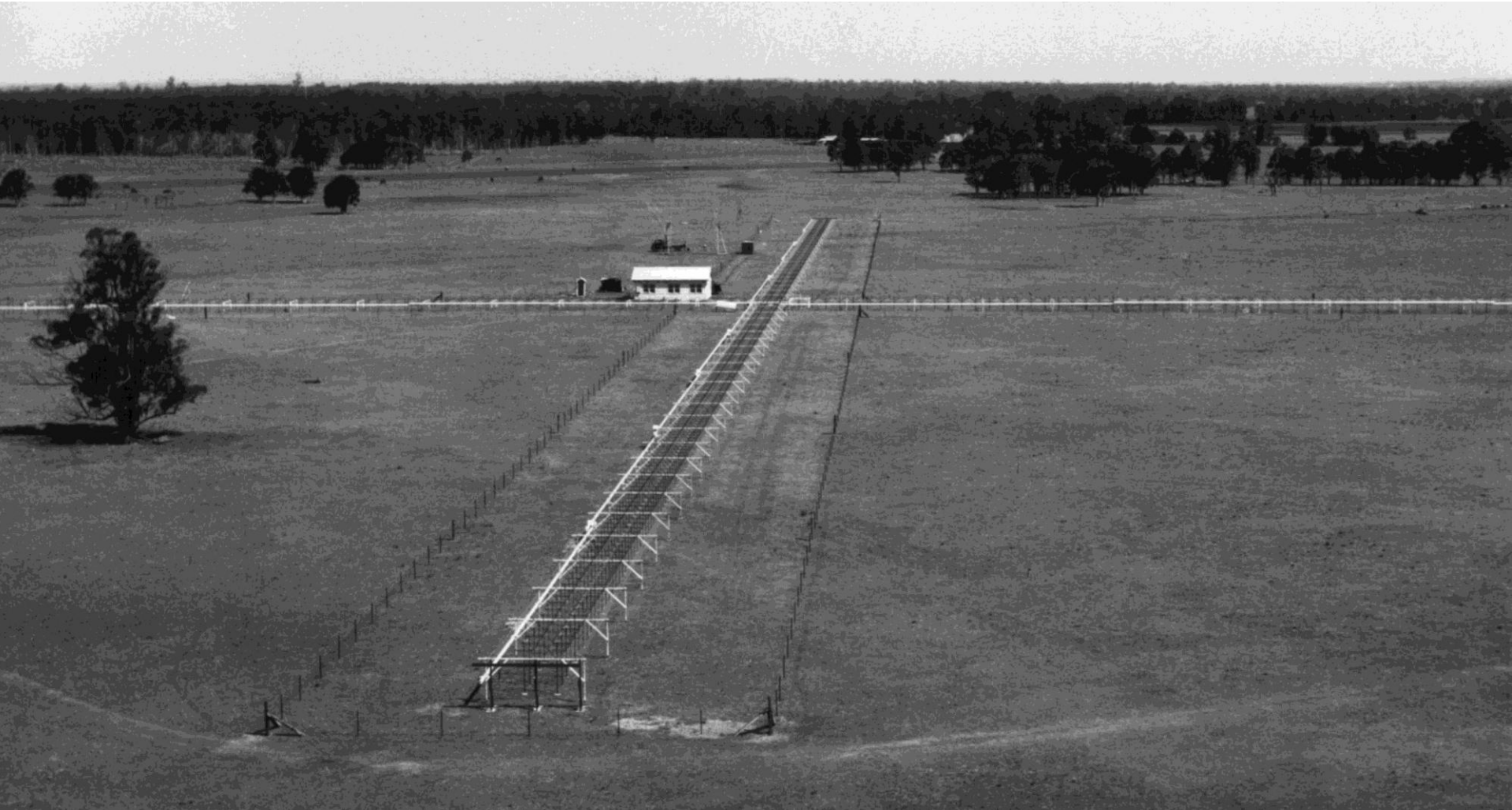
Bernie Mills

Alec Little

Mills Cross at Fleurs

- Joined by Alec Little
- Constructed 1953-45
- 85.5 MHz Observations
48 arcmin beamwidth

Mills Cross at Fleurs



Southern Sky Survey - MSH

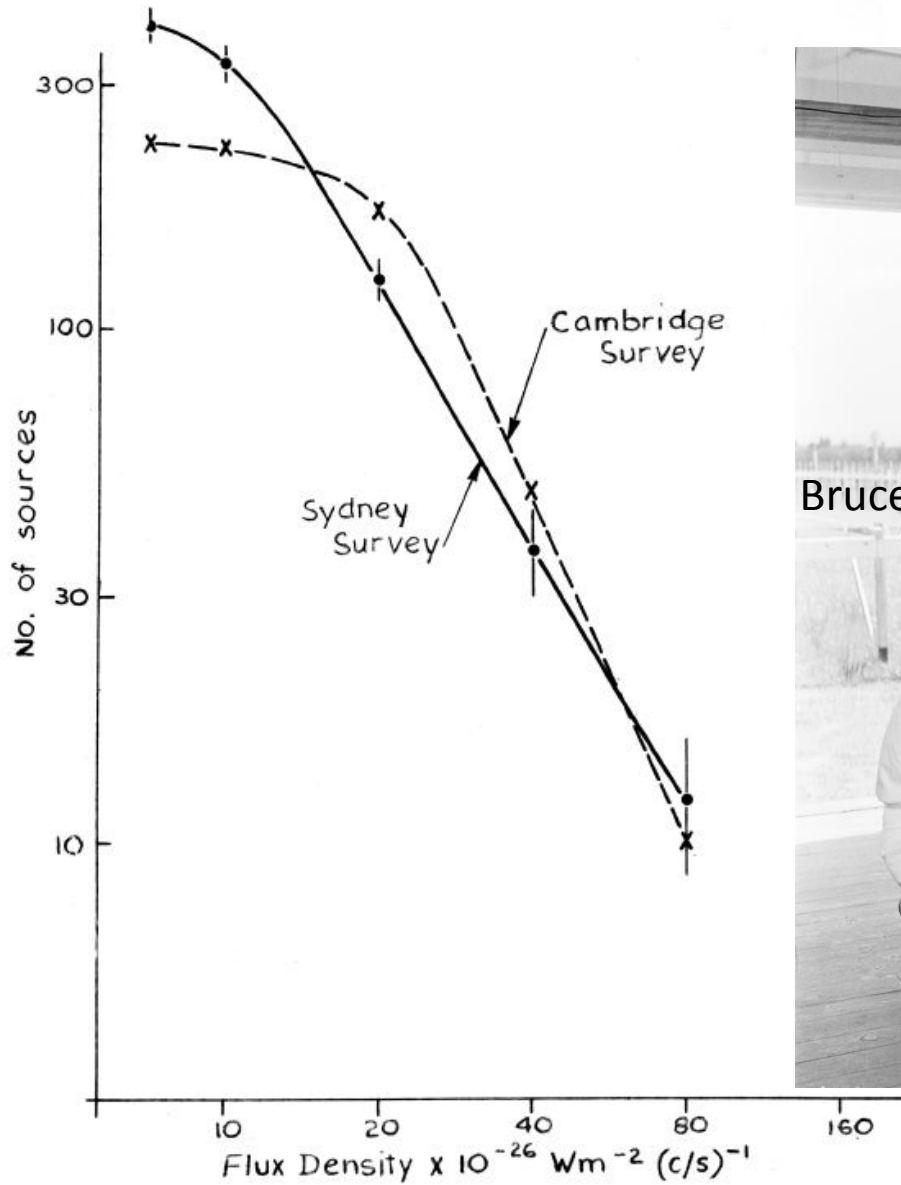
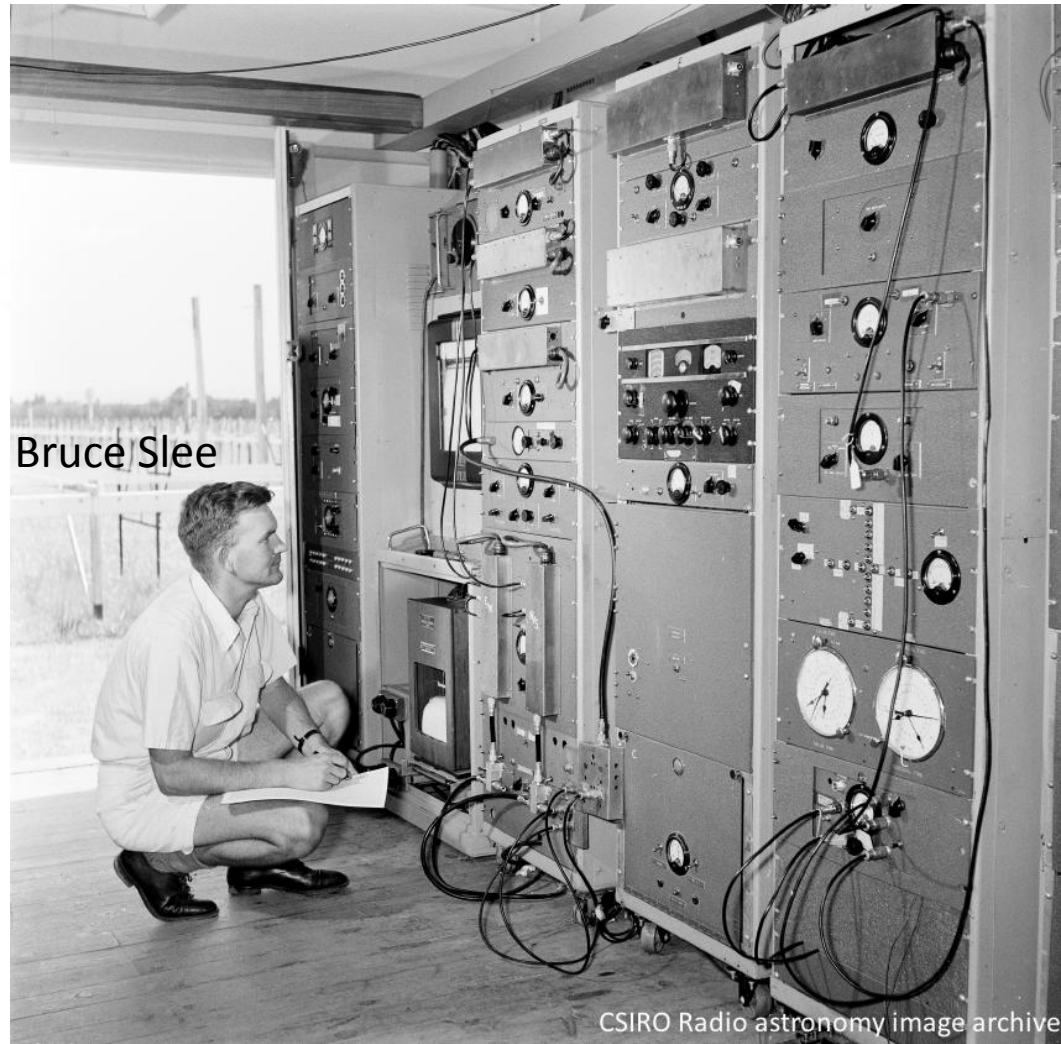
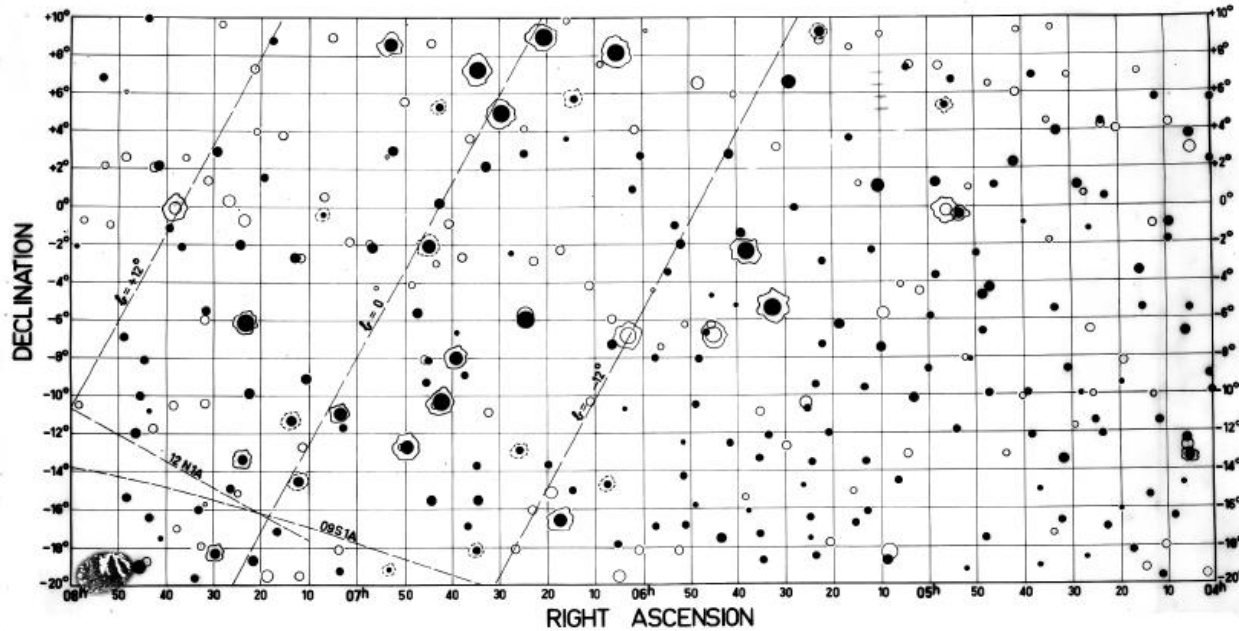


FIG. 5



Controversy with Cambridge



Response Patterns



Cambridge

Sydney

Flux density
 $\times 10^{26} \text{ Wm}^{-2} (\text{c/s})^{-1}$

> 80
 40-79.9
 20-39.9
 10-19.9
 < 10

Sydney Source

Cambridge Source

●	○
●	○
●	○
●	○
●	○

When sources are listed as "extended" or "large" in either catalogue, they are surrounded by an irregular line.

Conclusion

- In 1958 Mills attended the Paris Symposium on Radio Astronomy and then IAU General Assembly in Moscow
- Awarded Doctor of Science in Engineering for thesis covering development of the Mills Cross
- End of an Era → Schism at CSIRO Radiophysics



Bernie Mills

Chris Christiansen