Reflections on the discovery of pulsars (pulsating radio stars)

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Topics

- A story from radio astronomy the discovery of pulsating radio stars (pulsars)
- Pulsars today
- Pulsars nearly discovered: some 'near miss' stories
- Factors that led to the discovery

On arrival in Cambridge.....



Radio astronomy grad students presented with a set of tools

First build your radio telescope (through hail, rain and sunshine)



Typical working conditions for a PhD student

2048 81.5MHz λ/2 antennae (16 E-W rows of 64 + 64), 1000+ wooden posts, 120 miles/192 km wire and cable, area 57 tennis courts. Grant £12k. 6 people for 2 years.

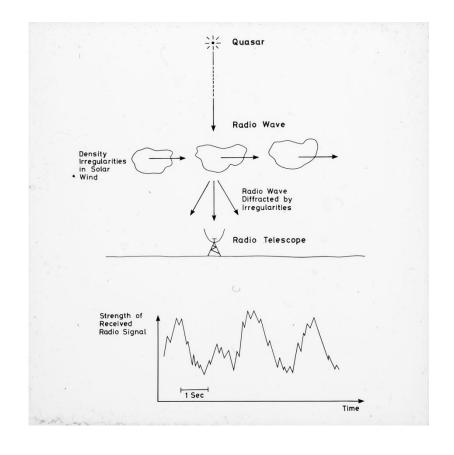
The 4.5 acre* radio telescope (*1.8 hectares), looking W



Phased interferometric array. Valve receiver!

Interplanetary scintillation

- Compact radio source scintillates, extended does not.
- Quasars (quasi-stellar radio sources) are compact; other radio sources are not



Discovery of pulsars

- Repeated mapping of the sky to identify quasars (and measure their angular diameter)
- Short time constant (short integration time) to study rapidly varying phenomenon (interplanetary scintillation). $\tau = 0.1$ s.
- 6 months' observing, starting July 1967

Data analysis

No computer!

3-pen chart paper

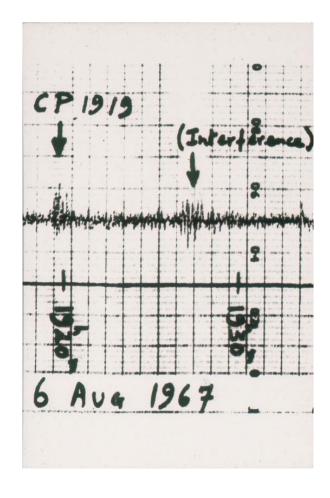
- 100' (30m) / day
- 400' (120m)/sky scan
- 3.3 miles (5.3km) total



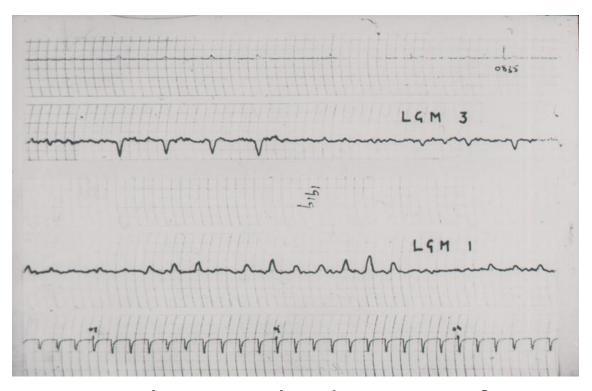
Discovery of pulsars II

Occasionally ¼"

 (0.5 cm) in the 400'
 (120 m) showed an unusual signal.



High-speed recording



Centre trace shows pulsed nature of emission

What is it?

- Local radio interference? No keeps sidereal time.
- Faulty equipment? No seen by a separate telescope and receiver.
- Small (short pulses) and big (maintains pulse period accurately). ??????

What is it (contd)?

- Dispersion measurement showed it was 200 light years away (i.e. beyond the Solar System, but nearby in the Milky Way)
- Little Green Men? They'd be on a planet orbiting their Sun; no Doppler effect
- Finding second, third and fourth

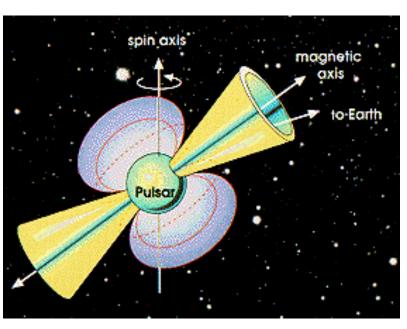
The naming of pulsars

- Interviewed by Science Correspondent of The Daily Telegraph – Anthony Michaelis – shortly after the discovery
- What were we going to call them?
- He suggested pulsar cf quasar

The Daily Telegraph

Pulsars today

How we picture pulsars today



Sound of a typical pulsar 11 revs per sec



Sound of a fast pulsar 700 revs per sec



Today.....

- About 2000 pulsars known, mostly in near half of our Galaxy
- Spinning, compact stars.
- Pulsar period = period of rotation of star
- 1.39ms < P < 8s; B ~ 10^5 10^{10} T
- Neutron (rich) stars: R = 12 km
- $M = 1.4 2.0M_{Sun}$, $(1M_{Sun} = 2 \times 10^{27} \text{ tonnes})$

Today contd

- 10⁻¹² > dP/dt > 10⁻²¹; the period of a typical pulsar has increased by about 1 second since the age of the dinosaurs!
- Serve as clocks for experimental relativity
- Few x 10⁵ active PSRs in Galaxy



Today.....

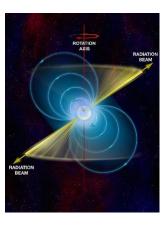
- Broad-band, but steep spectrum radio emitters. Few optical/X-ray emitters.
- Fermi satellite discovering gamma-ray pulsars
- HESS Cerenkov gamma ray array finding pulsar wind nebulae (like Crab nebula).



Today contd.

- Field still has not stabilised significant and surprising results still rolling in
- Unresolved issues:
 - Rotational energy loss, radiation mechanism, wind
 - Population (magnetars, intermittent...)
 - B, velocity, mass
 - Evolution

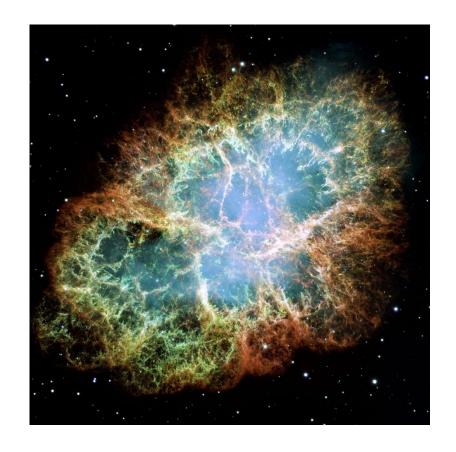
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Pulsars *nearly* discovered!

Crab Pulsar almost discovered...

 Bell and Hewish measurement of the angular diameter of the compact source in the Crab nebula. Using IPS. Integration time 0.1s (P=.033s). March 1967 paper.

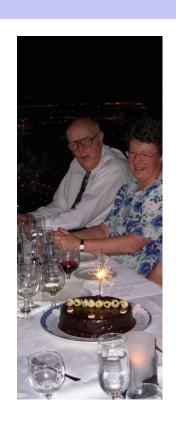


Crab PSR almost discovered II

- Late summer 1957; an Open Night at McDonald; 82 inch (Struve) telescope trained on Minkowski's star. Elliott Moore (newly graduated from Chicago) was
 - assisting.
- No instruments existed to provide follow-up.

Crab Pulsar – almost - III

- E-mail message received by ATNF, June 2007 from Charles Schisler:
- "As a USAF technician during a one year period back in 1967-68 at a Ballistic Missile Early Warning Site in Alaska I discovered fourteen pulsing signals on our extremely powerful radars...."



Crab PSR almost discovered IV

- Sue Simkin (NRAO) Oct/Nov 1967
- Carnegie Image Spectrograph, Kitt Peak 84-inch telescope.
- Lo Woltjer asked her to take a spectrum of Minkowski's star.
- Spectrum dull, but Sue observed flickering, or as if there were waves going out from it.
- •LW said it couldn't be, but when PSR discovered said it must have been.

PSR 0329+54 almost discovered

- 408 MHz survey (Europe) using large telescopes
- Last week of observing;
 one pen-recorder misbehaving
- Early hours of morning, that pen recorder started rhythmically sweeping
- Observer said 'Damn!' and thumped pen recorder – behaviour stopped!



PSR0329 +54 contd

- Observer said 'Good!', put on his coat and went home!...... He had made the first observation of PSR 0329+54.
- No entry in log book
- Following the discovery of the first pulsars (which did not include 0329+54) no search of their data for pulsars.

Tribute to Franco Pacini - insight

- A few months before discovery of pulsars announced Pacini published paper in Nature
- On how Crab Nebula energized by a neutron star with large magnetic field, inclined to spin axis.
- Magnetic dipole radiation at rotation period cannot escape plasma; energy and momentum dumped in nebula.



Factors that lead to discovery

Key factors in the discovery

- Our own telescope and receiver I understood its behaviour
- As a research student I had time/space to follow-up anomalies; imposter syndrome
- One of the first observations with a short time constant (new area of phase space)
- Not aware of other observers likely to make and announce the discovery before us

Key factors, contd

- We had a good address a reputable laboratory.
- This was not one of the objectives, or targets of the programme
- If we had computerised the search, would the pulsars have been discovered?
- Low frequencies and the time domain have long been neglected!

The End

